

# Study of neutrino oscillations in the long base-line experiment K2K

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XL PNPI Winter School, 21 February 2006

# Outline

Phenomenology of neutrino oscillations

**K2K**

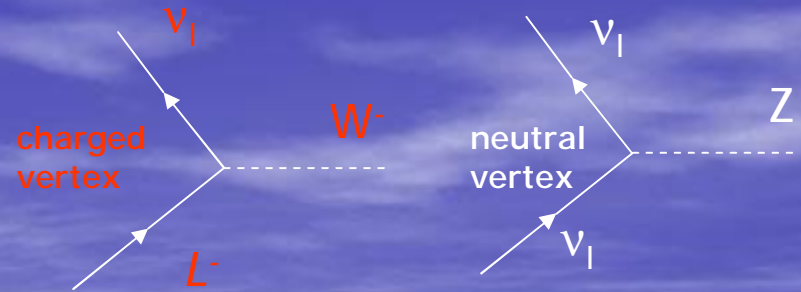
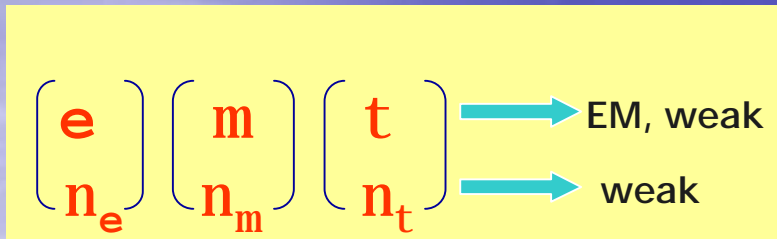
- principles of the experiment
- detectors
- performance
- analysis
- results

Neutrino oscillations: status and problems

**T2K**

Near future

# Leptons



two different neutrino bases

Flavor states	$n_e$	$n_m$	$n_t$
produced via charged currents	$L_e=+1$	$L_\mu=+1$	$L_\tau=+1$
Mass eigenvalues	$m_i$		
Mass eigenstates	$ n_i\rangle$		

flavor eigenstates  $|n_a\rangle$   $\stackrel{1}{\longleftrightarrow}$  mass eigenstates  $|n_i\rangle$

Unitary transformation:  $|n_a\rangle = U_{ai} |n_i\rangle$   
 $U_{ai}$  - neutrino(lepton) mixing matrix

# Mixing in two families

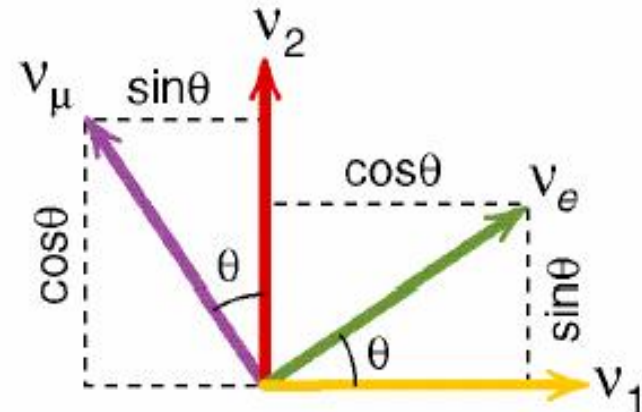
Consider for simplicity two families. Mixing matrix depends of a single parameter, the mixing angle  $\theta$

The weak and mass eigenstates are connected by a simple two-dimensional rotation

$$\begin{pmatrix} n_e \\ n_m \end{pmatrix} = \begin{pmatrix} \cos q & \sin q \\ -\sin q & \cos q \end{pmatrix} \begin{pmatrix} n_1 \\ n_2 \end{pmatrix} = U \begin{pmatrix} n_1 \\ n_2 \end{pmatrix}$$

$$n_e = \cos q |n_1\rangle + \sin q |n_2\rangle$$

$$n_m = -\sin q |n_1\rangle + \cos q |n_2\rangle$$

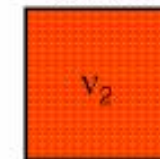


Mass states

First

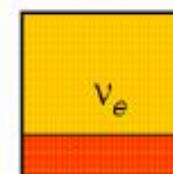


Second

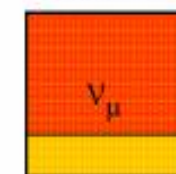


Weak states

First



Second



# Neutrino oscillations



The weak interaction produces neutrinos of a given flavor

Distance  $x_0 = 0$  time  $t=0$

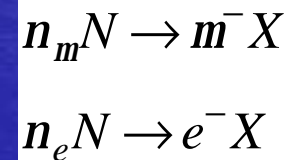
$$|n(x_0)\rangle = |n_e\rangle = \cos q |n_1\rangle + \sin q |n_2\rangle$$

The mass eigenstates propagate at different velocities

Distance  $x$  time  $t$

$$|n_e\rangle_t = \cos q \exp(ip_1 x) \exp(-iE_1 t) |n_1\rangle + \sin q \exp(ip_2 x) \exp(-iE_2 t) |n_2\rangle$$

Detection via weak interactions



$$P(n_e \rightarrow n_m) = |\langle n_m | n(t) \rangle|^2$$

# Oscillation formalism

$$E^2 = p^2 + m^2 \quad \text{neutrino: } p \gg m \Rightarrow E \approx p + \frac{m^2}{2p} \quad E_2 - E_1 = \frac{\Delta m^2 L}{2E}$$

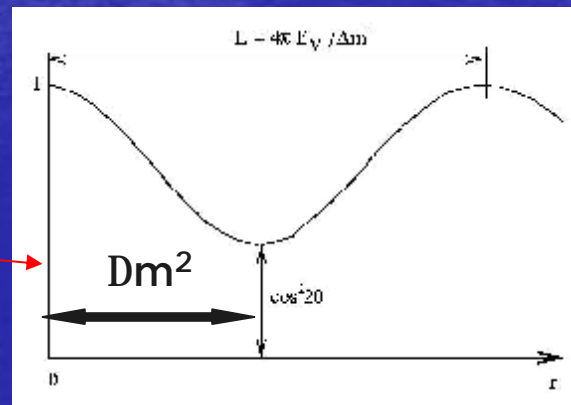
$$\Delta m^2 = |m_1^2 - m_2^2| \quad E \gg p$$

$$P(n_e \rightarrow n_m) = |\langle n_m | n(L) \rangle|^2 = \left| -s c e^{-i \frac{m_1^2}{2E} L} + c s e^{-i \frac{m_2^2}{2E} L} \right|^2$$

$$= 4s^2 c^2 \left( 1 - \cos \frac{m_1^2 - m_2^2}{2E} L \right) = \sin^2(2q) \sin^2 \left( \frac{\Delta m_{12}^2}{4E} L \right)$$

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - P(\nu_\mu \rightarrow \nu_\tau)$$

$P(\nu_l \rightarrow \nu_l)$



$$P(n_m \textcircled{R} n_x) = \sin^2 2q \sin^2 [1.27 \Delta m^2 (\text{eV}^2) L (\text{km}) / E_n (\text{GeV})]$$

$$P(n_m \textcircled{R} n_m) = 1 - \sin^2 2q \sin^2 [1.27 \Delta m^2 (\text{eV}^2) L (\text{km}) / E_n (\text{GeV})]$$

# PMNS mixing matrix

3 families

$$\begin{pmatrix} n_e \\ n_m \\ n_t \end{pmatrix} = U \begin{pmatrix} n_1 \\ n_2 \\ n_3 \end{pmatrix} \quad U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{m1} & U_{m2} & U_{m3} \\ U_{t1} & U_{t2} & U_{t3} \end{pmatrix}$$

**$U$  parameterization**

three mixing angles  $\theta_{12}$   $\theta_{13}$   $\theta_{23}$   
 complex phase  $\delta$

$$\frac{|U_{e2}|^2}{|U_{e1}|^2} = \tan^2 \theta_{12} \quad \frac{|U_{m3}|^2}{|U_{t3}|^2} = \tan^2 \theta_{23} \quad U_{e3} = \sin \theta_{13} e^{-i\delta}$$

$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$

$$\Delta m_{12}^2 + \Delta m_{23}^2 + \Delta m_{31}^2 = 0$$



two independent  $\Delta m^2$

$$\Delta m_{12}^2 = \Delta m_{sol}^2 \quad \Delta m_{23}^2 \cong \Delta m_{31}^2 = \Delta m_{atm}^2$$

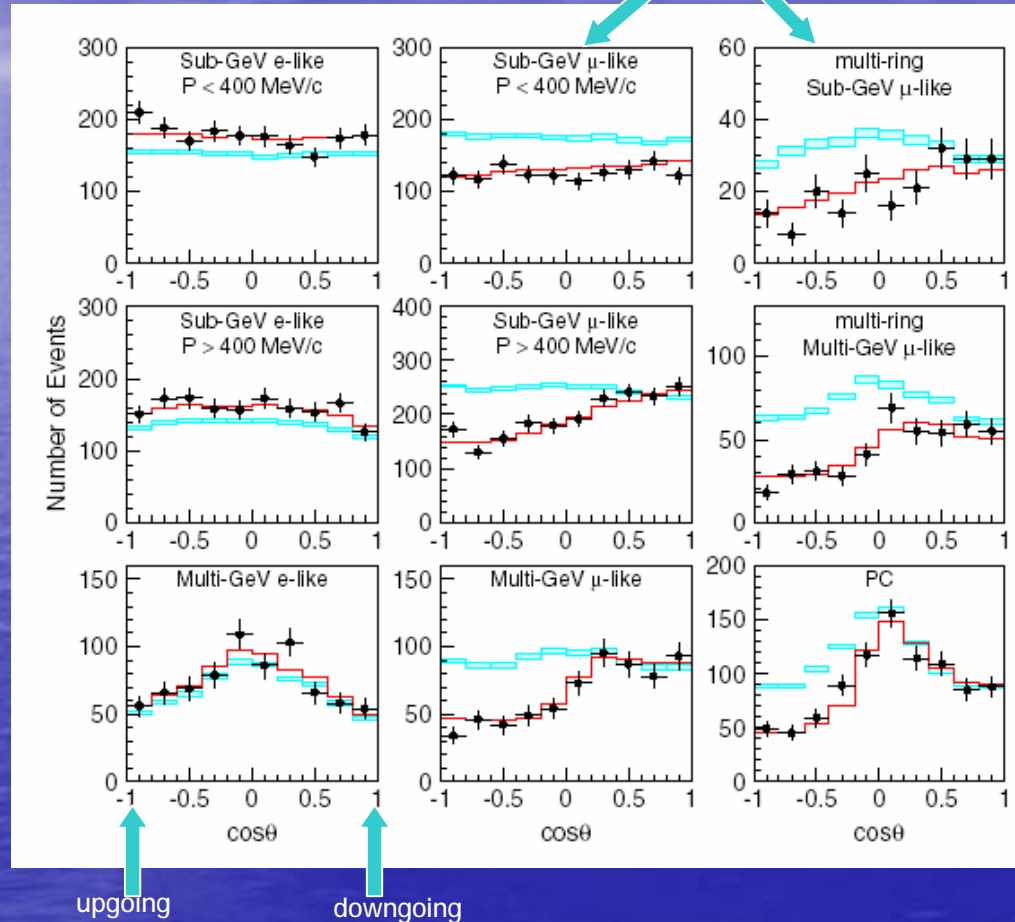
# Atmospheric neutrino oscillations

$0.2 < \cos\theta < 1$      $L = 20\text{-}500 \text{ km}$   
 $-1 < \cos\theta < -0.2$      $L = 500\text{-}13000 \text{ km}$

$n_e$

$n_m$

SK →



$(U/D)_m = 0.54 \pm 0.04 \pm 0.01$

**SK:**  $n_m$  oscillation with  $\Delta m^2 \sim (2\text{-}3) \times 10^{-3} \text{ eV}^2$



# Main goal of K2K

First accelerator long base-line neutrino experiment

Measurement of (search for) neutrino oscillations in LBL accelerator experiment to confirm the oscillation observed by the SuperKamiokande

$$\Delta m^2 \sim (2-3) \times 10^{-3} \text{ eV}^2 \quad \sin 2\theta \sim 1$$

# Experiment K2K

Collaboration K2K: Japan-USA-Korea-Canada-Russia-France-Italy-Spain-Switzerland

$$n_m \otimes n_t \quad (n_m \otimes n_x)$$

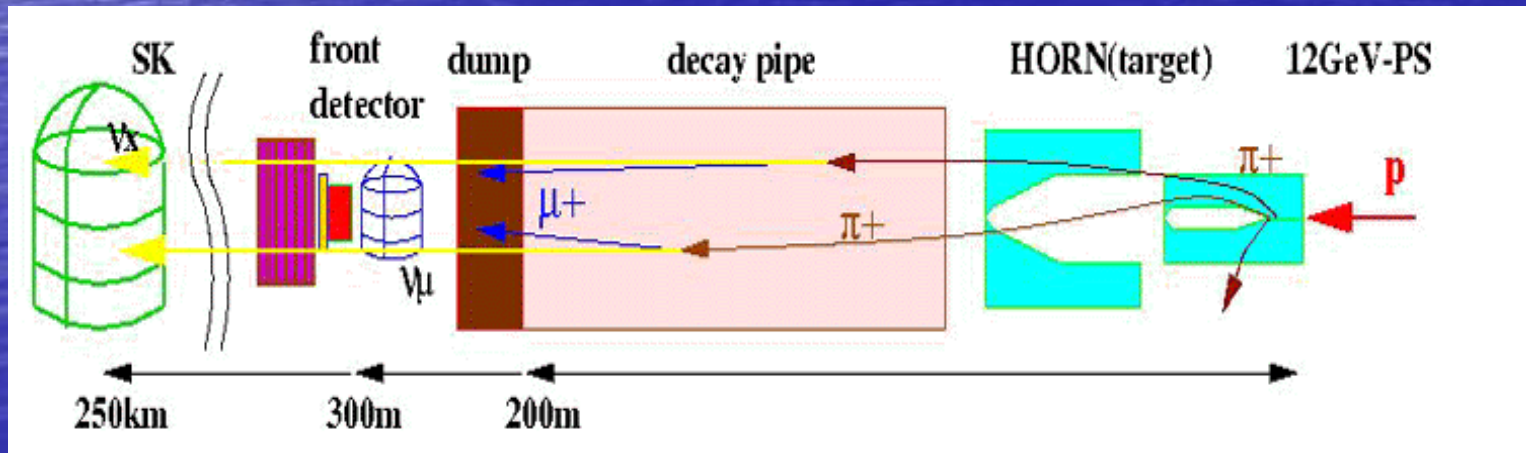
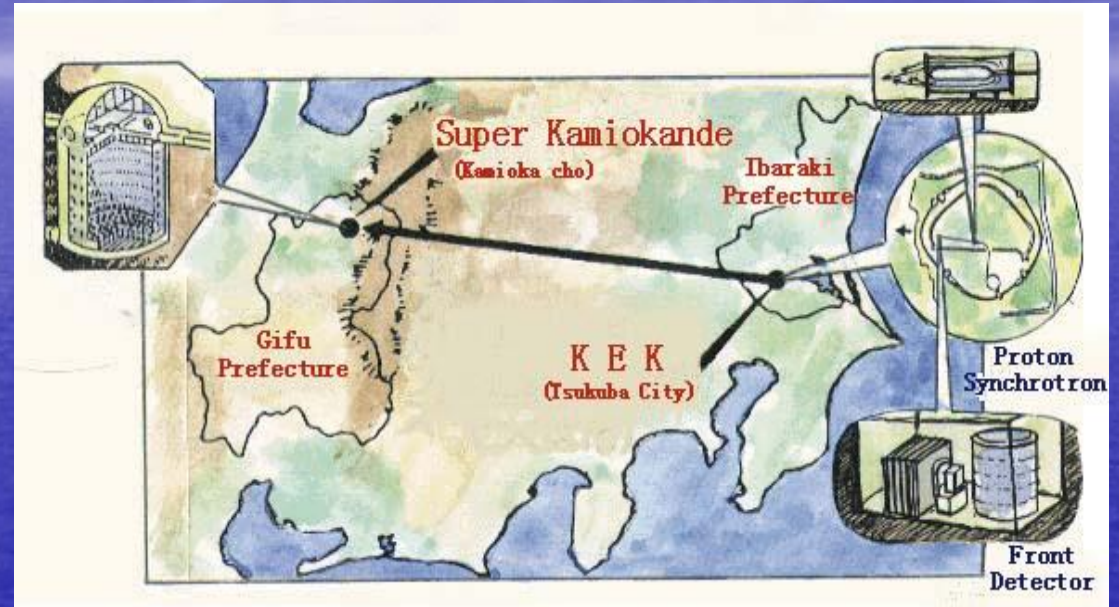
$$L/E_n \gg 200$$

$$L=250 \text{ km} \quad \langle E_n \rangle \gg 1.3 \text{ GeV}$$

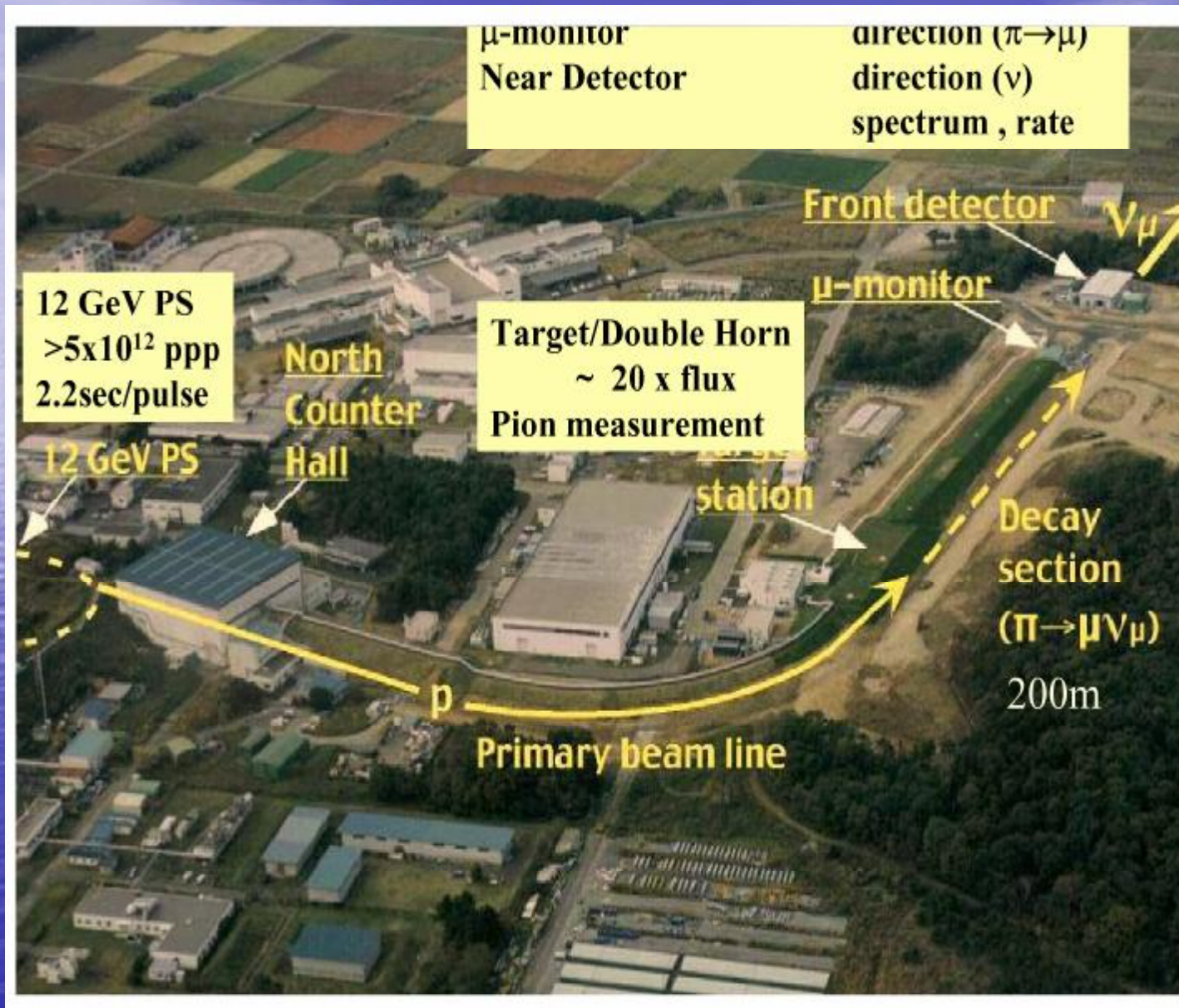
$$n_m \quad 98.2\%$$

$$n_e \quad 1.3\%$$

$$Dm^2 \approx 2 \cdot 10^{-3} \text{ eV}^2$$

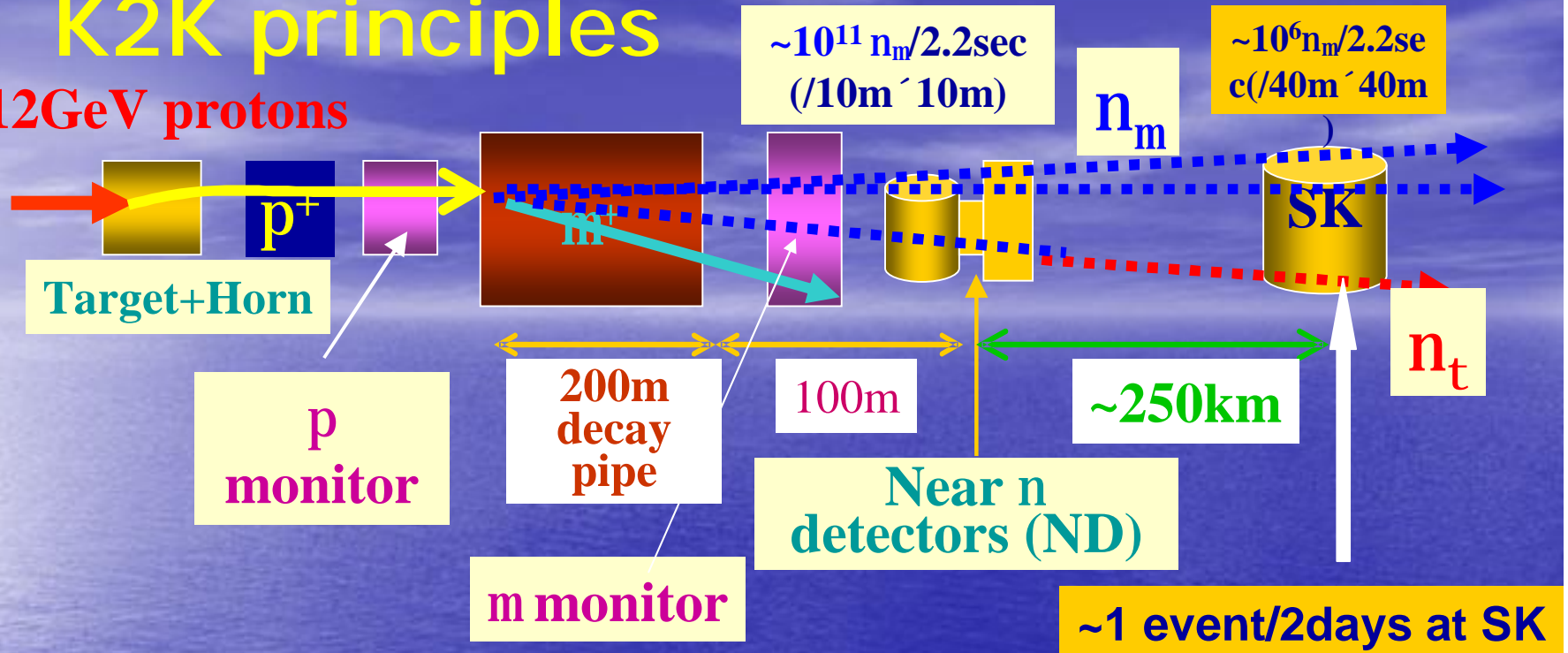


# K2K Beam Line



# K2K principles

12GeV protons



$$n_m \text{ flux at SK} = n_m (\text{flux at ND}) \times (\text{Far/Near})$$

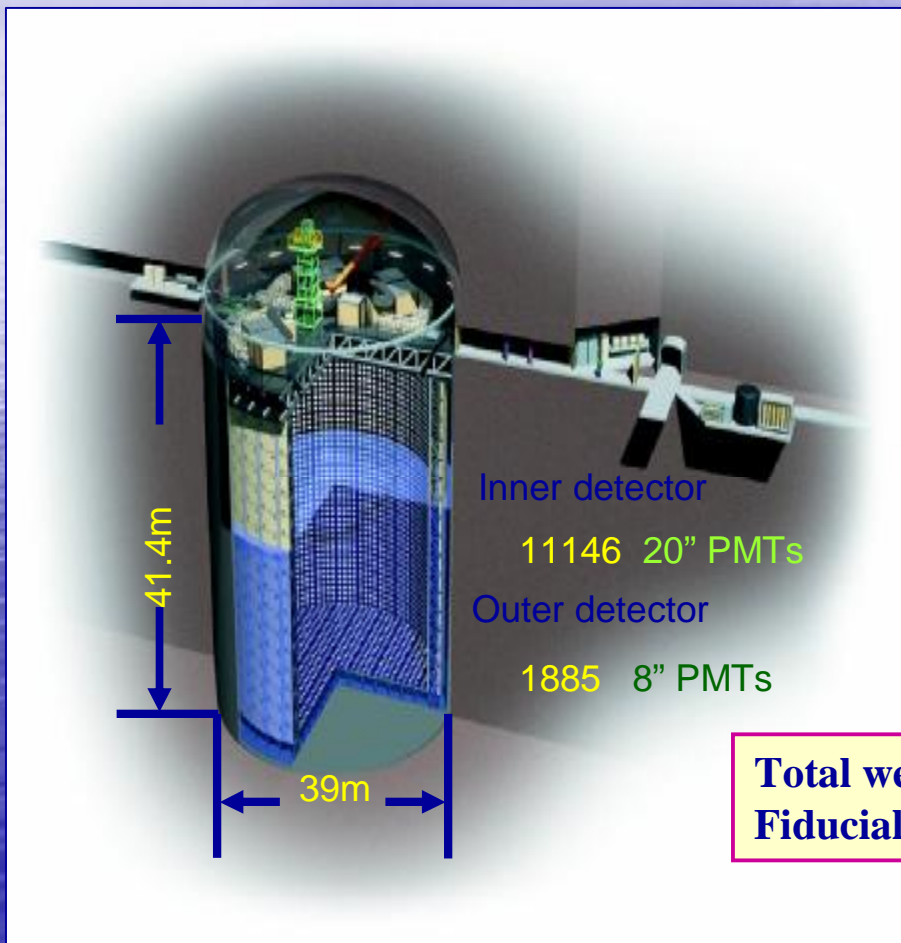
Far/Near ratio (by MC)  $\sim 10^{-6}$

## Signal of $n$ oscillation

- | Reduction of  $n_m$  events
- | Distortion of  $n_m$  energy spectrum

# Far detector

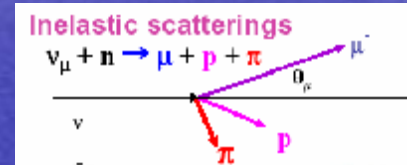
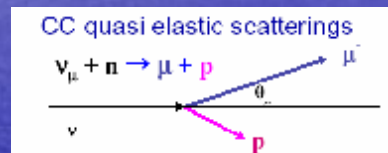
## Super-Kamiokande I



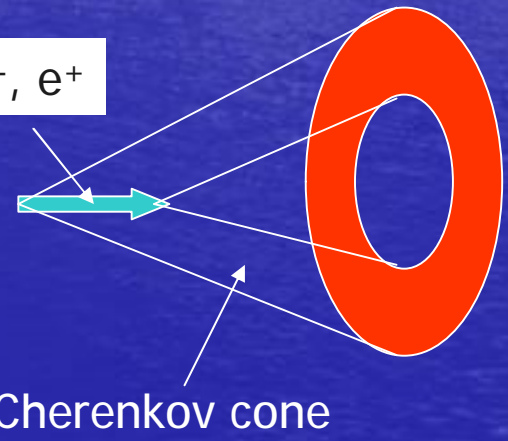
**Total weight 50 kt**  
**Fiducial 22.5 kt**

## Super-Kamiokande II

~5200 PMTs  
with FRP+Acrylic cover



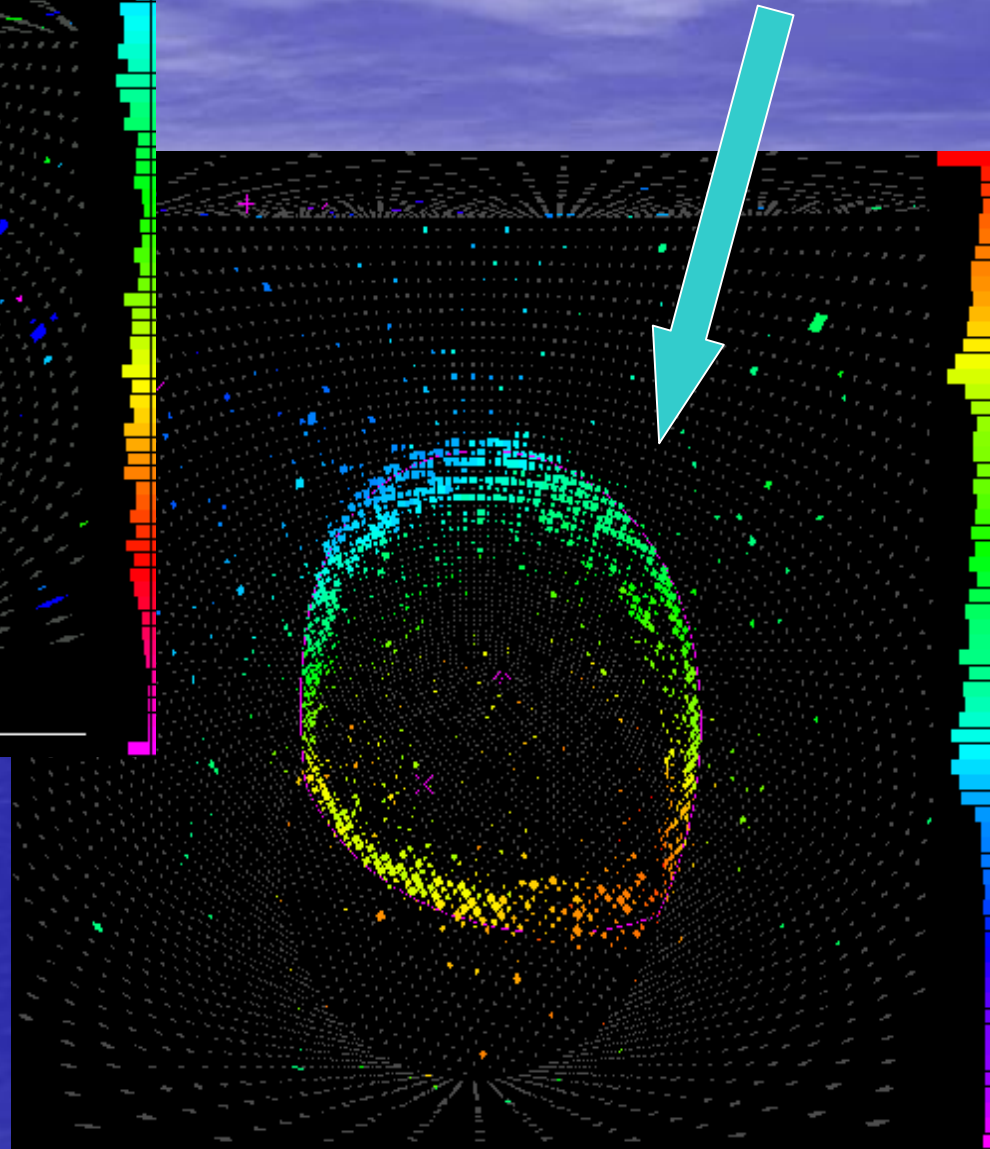
$\mu^{-}, e^{-}, e^{+}$



# SK events

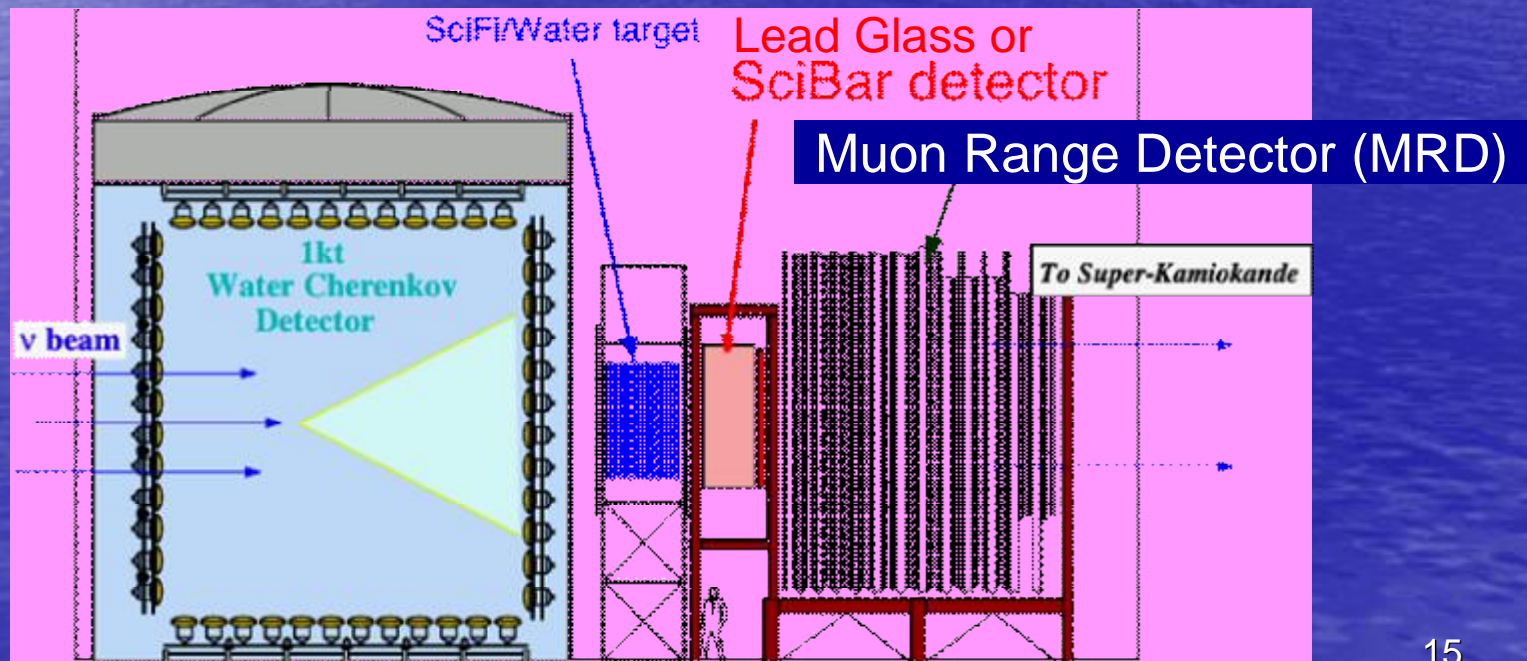
m - like

e - like

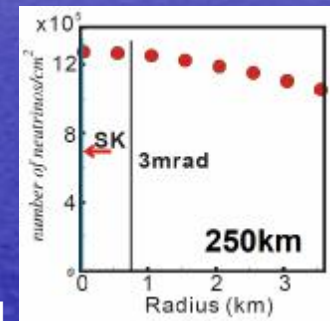
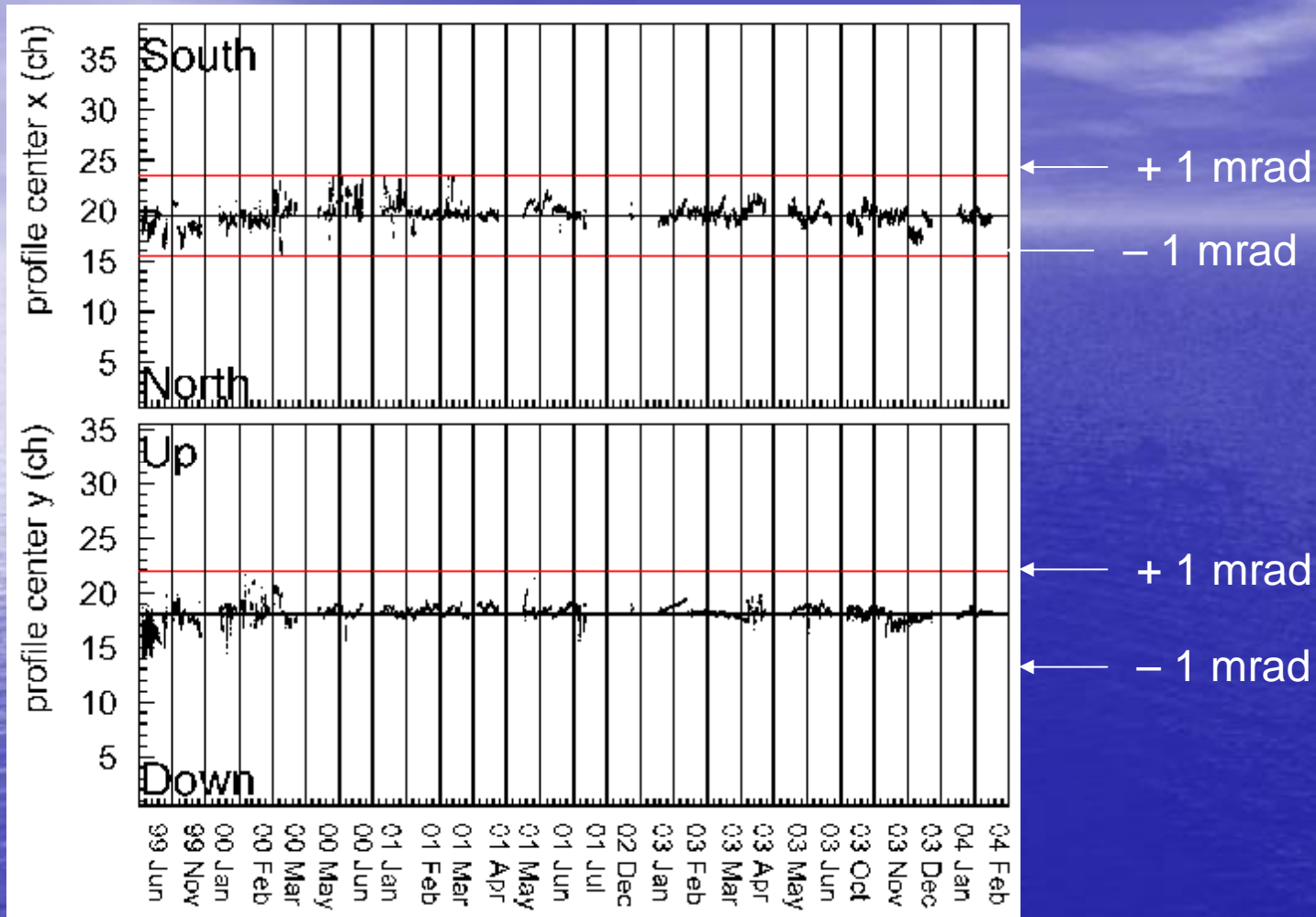


# Near Detectors

- **1KT**: water Cherenkov detector [25t fiducial]
- **SciFi**: scintillating fiber and water target [6t fiducial]
- **LG**: Lead glass calorimeter (removed in 2002)
- **SciBar**: fully-active scintillator detector [10t fiducial] (installed in 2003)
- **MRD**: muon range detector



# Beam stability (muon monitor)

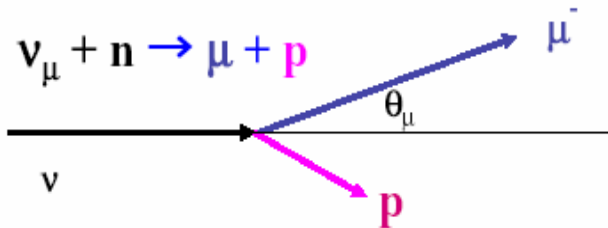


- muon profile is monitored spill-by-spill
- muon center is stable within 1mrad.

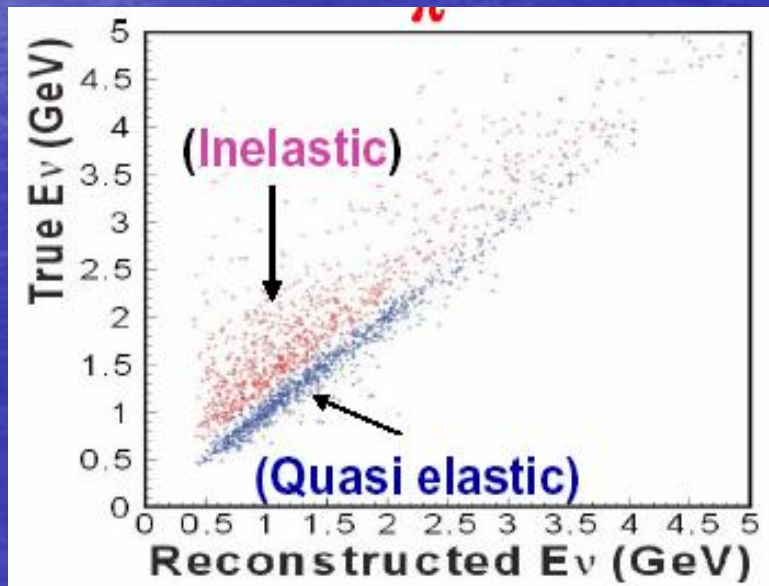
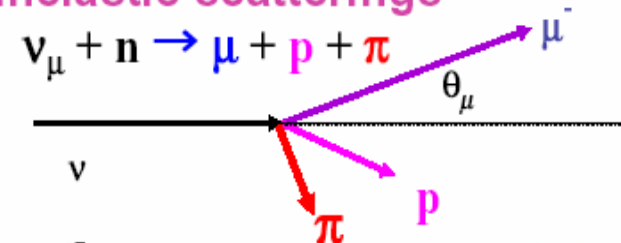


# Neutrino energy reconstruction

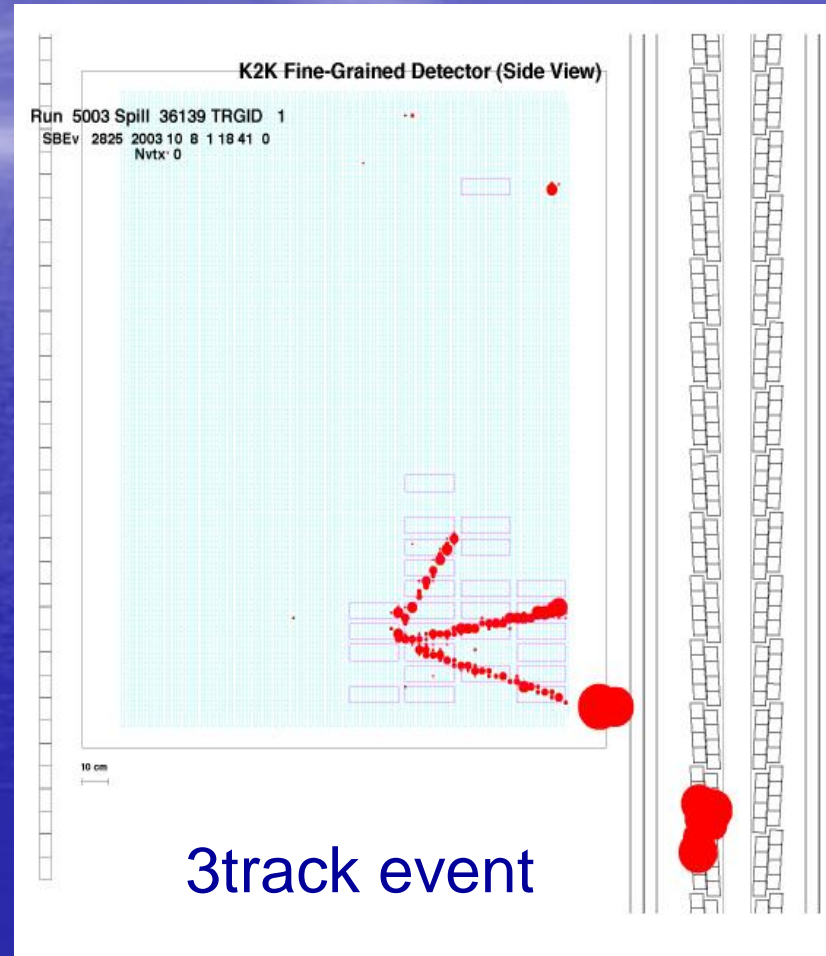
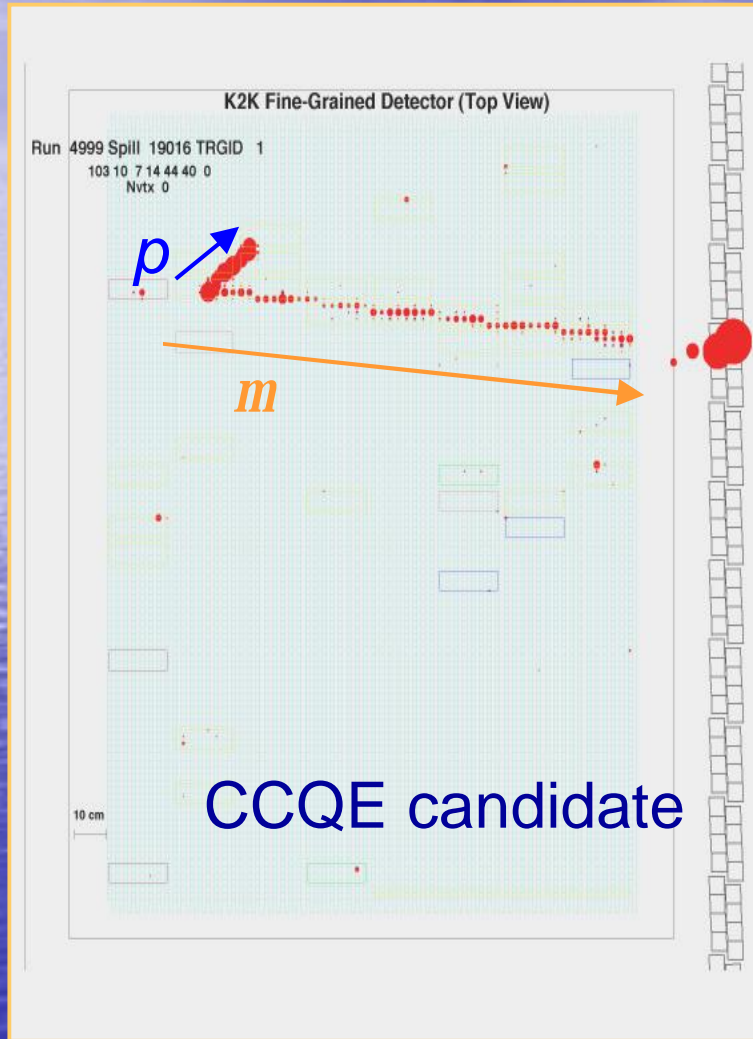
CC quasi elastic scatterings



Inelastic scatterings



# SciBar Event Display



# Measurements at ND

- 1KT :  $P_\mu < 1.5 \text{ GeV}/c$ ,  $4\pi$  acceptance

- 1-ring  $\mu$ -like ( $1R_\mu$ ) fully contained in Fiducial volume 25 ton (FC) :

- SciFi :  $P_\mu > 0.55 \text{ GeV}/c$ ,  $\theta_\mu < 60 \text{ deg}$ . Fid volume 5.6 t

- 1-track  $\mu$ -like : QE-like ( $\Delta\theta_p < 25 \text{ deg}$ ) :

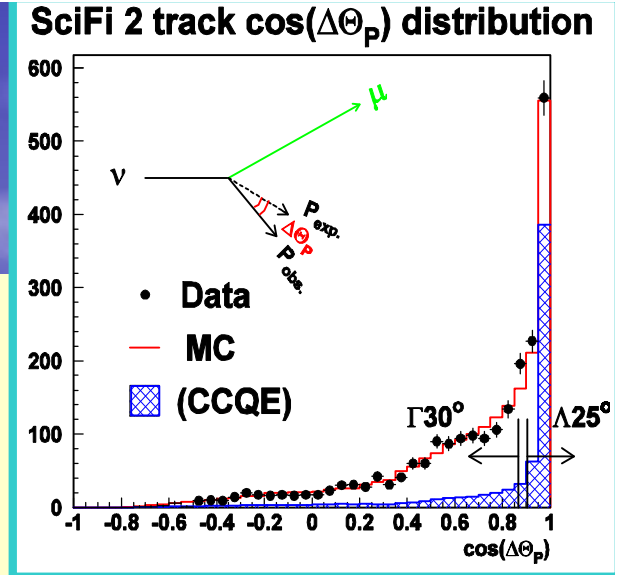
- nonQE-like ( $\Delta\theta_p > 30 \text{ deg}$ ) :

- Scibar :  $P_\mu > 0.45 \text{ GeV}/c$  Fid. volume 9.4 t

- PIMON

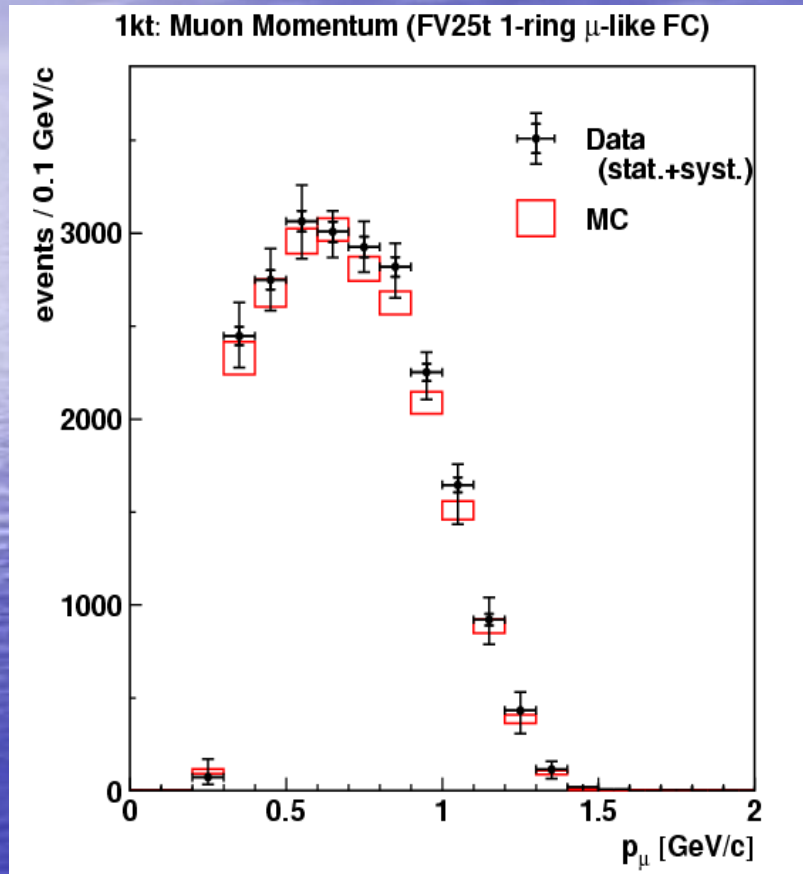
( $p, \theta$ ) distribution  $\Rightarrow$  Neutrino Spectrum ( $> 1 \text{ GeV}$ ) Fitting Parameters

$E_\nu$  : 8 bins, nonQE/QE ratio : 1

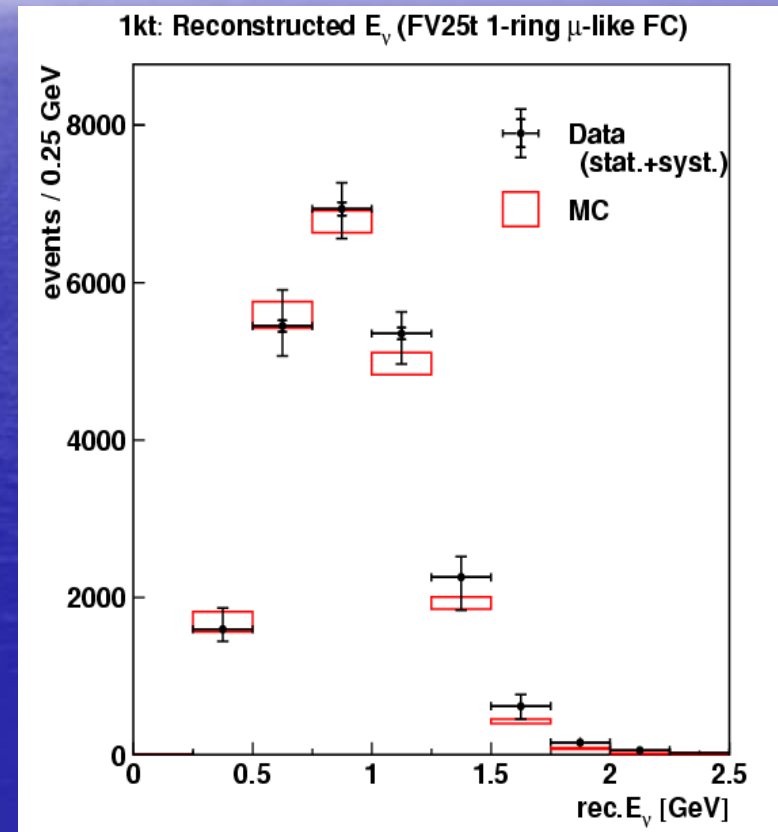


# 1kT spectra

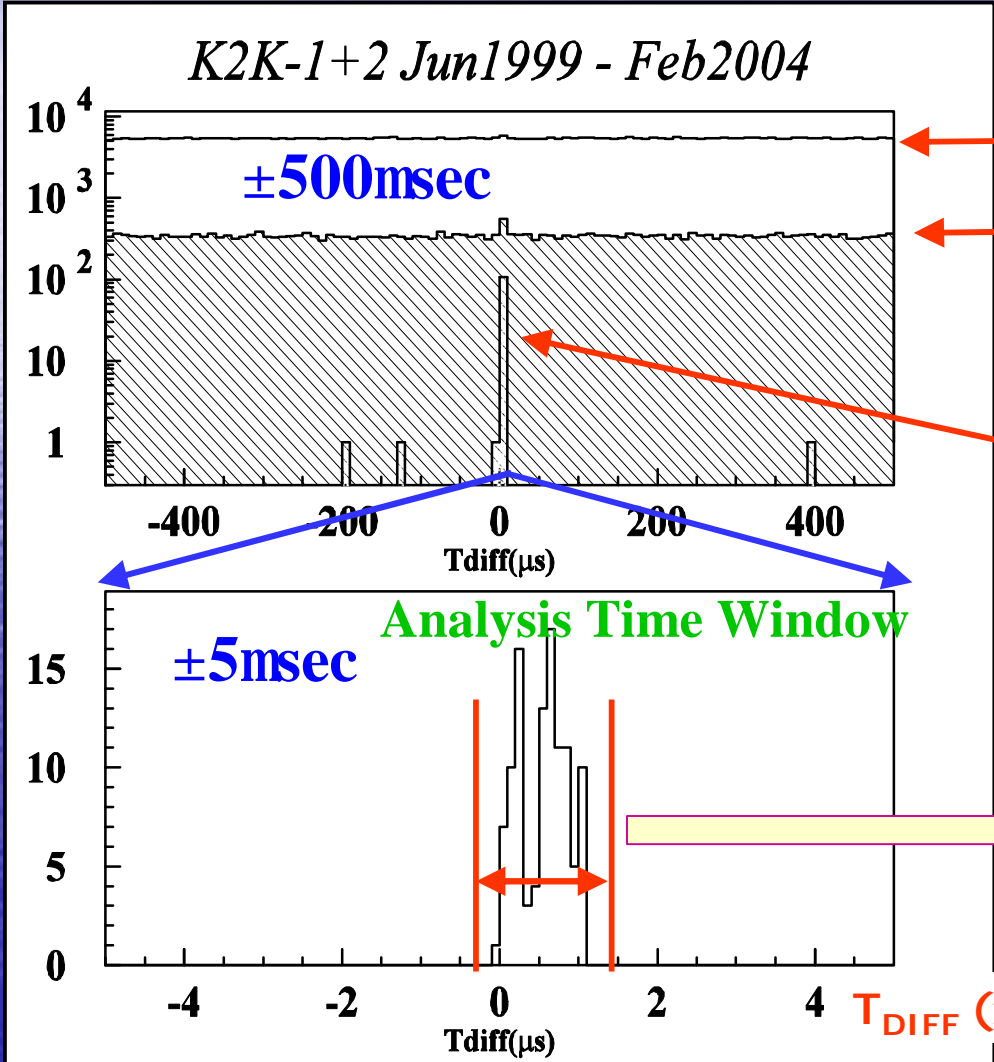
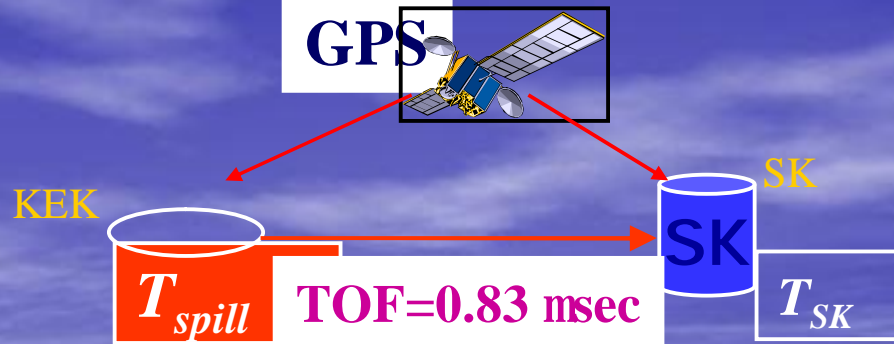
1 ring m-like



reconstructed neutrino energy



# n's from KEK



Decay electron cut

$\geq 20\text{MeV}$  Deposited Energy

No activity in Outer Detector (FV)  
 Event vertex in fiducial volume (FC)  
 More than 30MeV deposited energy

$$-0.2 < T_{SK} - T_{spill}^-$$

$$\text{TOF} < 1.3\text{msec}$$

(BG: 1.6 events within  $\pm 500\text{ms}$   
 $2.4 \times 10^{-3}$  events in 1.5ms)

# Oscillation analysis

- Total number of FCFV events
- $E_n$  spectrum shape of FCFV 1-ring muon events
- Systematic error term

$$L(Dm^2, \sin 2q, f^x) = L_{norm}(Dm^2, \sin 2q, f^x) \times L_{shape}(Dm^2, \sin 2q, f^x) \times L_{syst}(f^x)$$



Poisson probability  
for # FCFV events



Shape of  $E_n$  spectrum  
of 1-ring  $m$  events



Systematic error

$f^x$  - systematic error parameters

Normalization, Flux, and nQE/QE ratio are in  $f^x$

# Events at SK

Fully contained events in 22.5 kt fiducial volume of SK

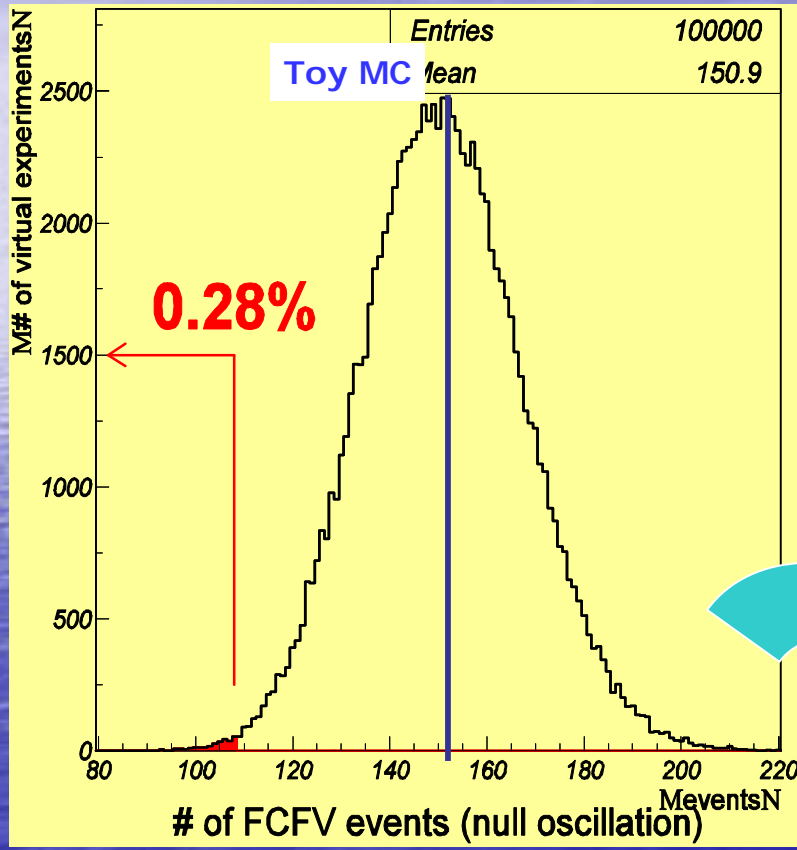
$$L_{norm}(\Delta m^2, \sin 2q, f^x)$$

$$L_{norm} = \frac{(N^{exp})^{N^{obs}}}{N^{obs}!} \cdot \exp(-N^{exp})$$

#SK Events



107



SK	Event summary	
	Data	MC
1-ring m-like	57	85.5
1-ring e-like	9	8.7
Multi-ring	40	56.7
<b>Total</b>	<b>107</b>	<b>150.9</b>

107 150.9

$$N_{SK}^{obs} = 107^{+12}_{-10} \quad N_{SK}^{exp} = 151$$

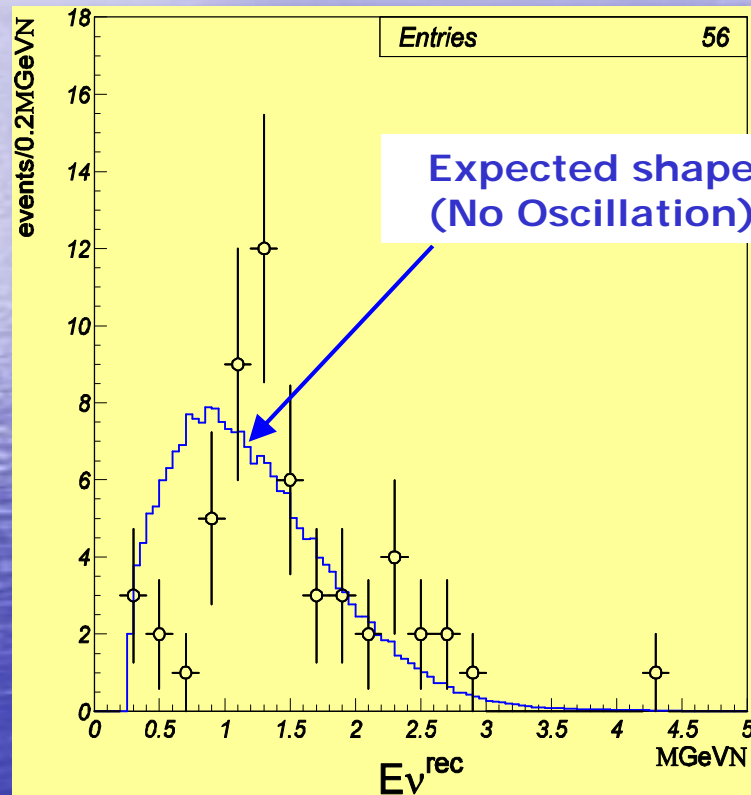
Only number of events  
No oscillation probability = 0.28%

# Shape analysis

$$L_{shape}(\Delta m^2, \sin 2q, f^x)$$

CC-QE assumption

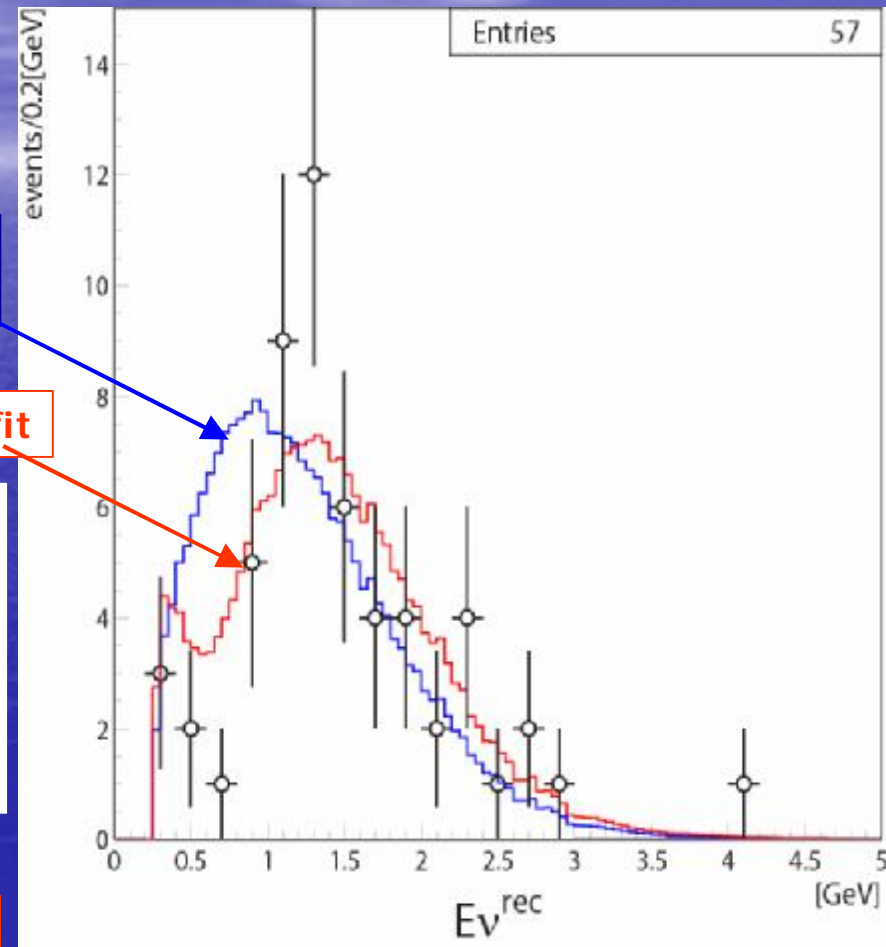
$$E_n^{rec} = \frac{(m_N - V)E_m - m_m^2/2 + m_N V - V^2/2}{(m_N - V) - E_m + p_m \cos q_m}$$



Only shape  
Kolmogorov-Smirnov test  
No oscillation probability = 0.74%



# Shape distortion



Expected shape  
(No Oscillation)

Best fit

Best fit value

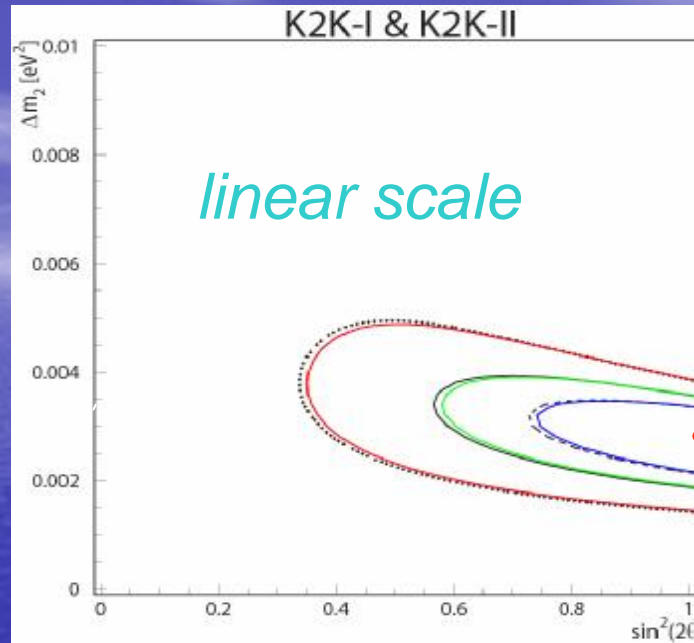
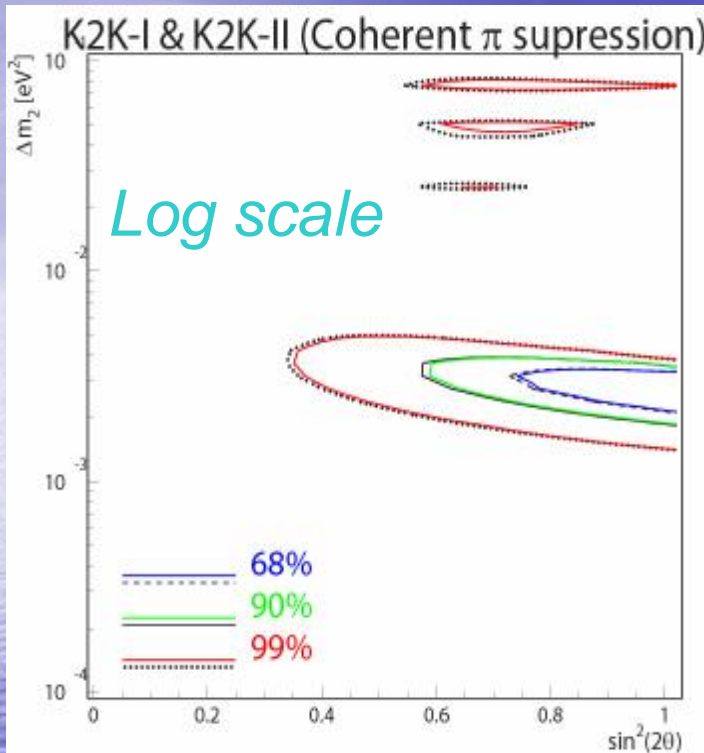
$$\sin^2 2\theta = 1.00$$

$$\Delta m^2 [\text{eV}^2] = (2.79 \pm 0.36) \times 10^{-3}$$

Kolmogorov-Smirnov test

Best fit probability = 36%

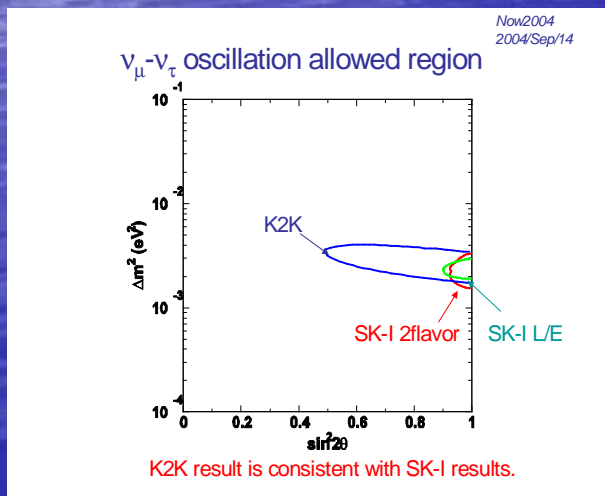
# Allowed region of oscillating parameters



Allowed region of  $\Delta m^2 @ \sin^2 2\theta = 1$   
 $1.99 \sim 3.33$  [eV<sup>2</sup>] (68%)  
 $1.9 \sim 3.6$  [eV<sup>2</sup>] (90%)

Null Oscillation probability  
 $0.0050\%$  ( $4.06\sigma$ )

Evidence of  $\nu_\mu$  oscillation in K2K experiment



# Search for $\nu_m \textcircled{R} \nu_e$ oscillation

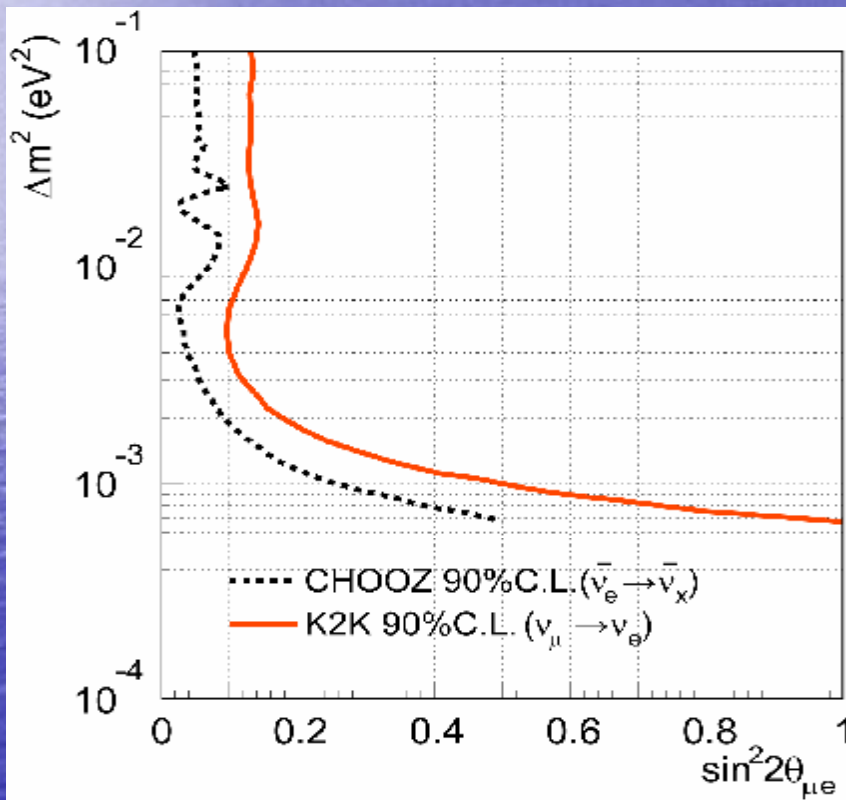
-K2K-1-	$\nu_\mu$ MC	beam $\nu_e$	Data
FCFV	81.1	0.81	55
Single ring	50.92	0.47	33
Electron like	2.66	0.40	3
Evis > 100 MeV	2.47	0.40	2
No decay-e	1.90	0.35	1
Pi0 cut	<b>0.58</b>	<b>0.17</b>	<b>0</b>

-K2K-2-	$\nu_\mu$ MC	beam $\nu$	Data
FCFV	77.4	0.86	57
Single ring	49.41	0.52	34
Electron like	3.21	0.44	5
Evis > 100 MeV	2.93	0.44	5
No decay-e	2.17	0.39	4
Pi0 cut	<b>0.74</b>	<b>0.21</b>	<b>1</b>

In total,  
 #expected BG = 1.70  
 #observed = 1

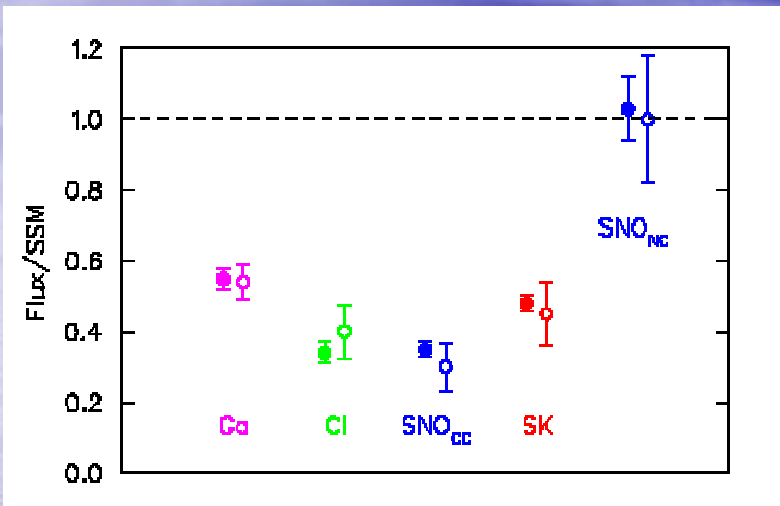
$$\bar{\nu}_m \dashrightarrow \bar{\nu}_e$$

$\Delta m^2$  vs.  $\sin^2 2\theta_{\mu e}$



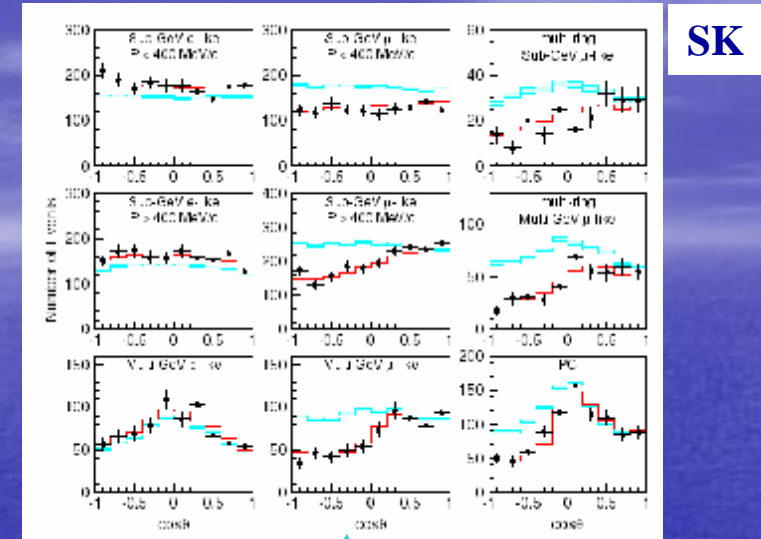
Assumption:  $2\sin^2 2\theta_{\mu e} = \sin^2 2\theta_{13}$

# Evidence of neutrino oscillations



solar

$$\Delta m^2 \sim (7-8) \times 10^{-5} \text{ eV}^2$$

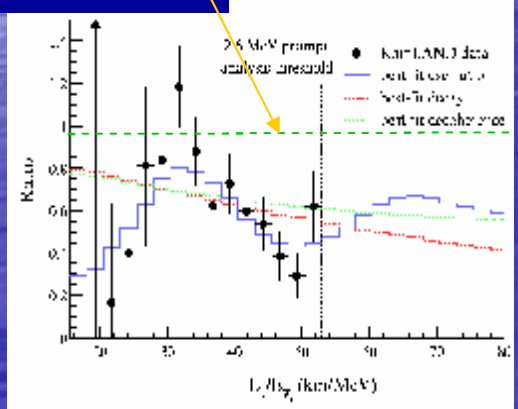


SK

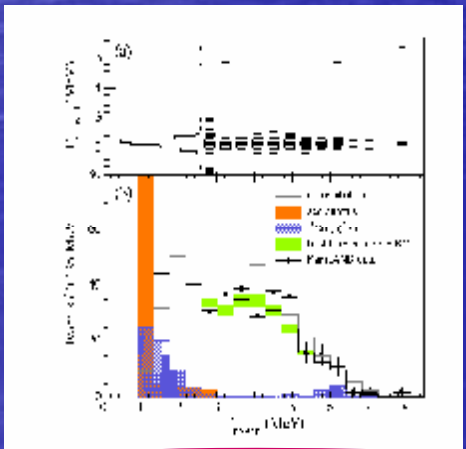
atmospheric

$$\Delta m^2 \sim (2-3) \times 10^{-3} \text{ eV}^2$$

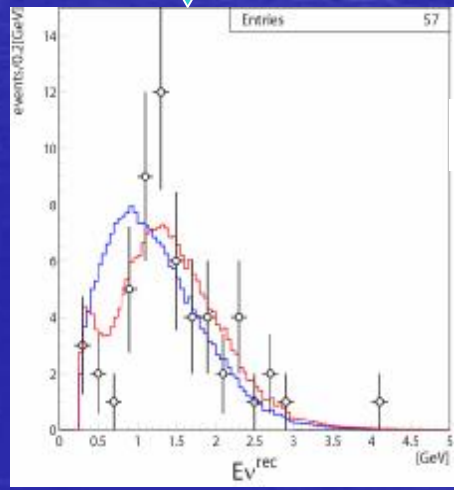
No oscillation



Kamland



$$\text{LSND} \rightarrow \Delta m^2 = 0.2 - 10 \text{ eV}^2 \rightarrow m_n > 0.4 \text{ eV}$$



K2K

# Neutrino masses and mixings

3 families

Oscillation parameters		
	central value	3 $\sigma$ interval
$Dm^2_{12}$ ( $10^{-5} \text{ eV}^2$ )	7.9	7.1 - 8.9
$Dm^2_{31}$ ( $10^{-3} \text{ eV}^2$ )	2.2	1.4 - 3.3
$\sin^2 \theta_{12}$	0.31	0.24 - 0.40
$\sin^2 \theta_{23}$	0.50	0.34 - 0.68
$\sin^2 \theta_{13}$	0.0	<0.047

LSND  $\Rightarrow Dm^2 = 0.2 - 10 \text{ eV}^2 \Rightarrow m_n > 0.4 \text{ eV}$

Mixing	Quarks	Leptons
1-2 $\theta_{12}$	$13^\circ$	$33^\circ$
2-3 $\theta_{23}$	$2.3^\circ$	$45^\circ$
1-3 $\theta_{13}$	$\sim 0.5^\circ$	$< 13^\circ$

## Challenges in neutrino physics

LBL accelerator experiments

reactor

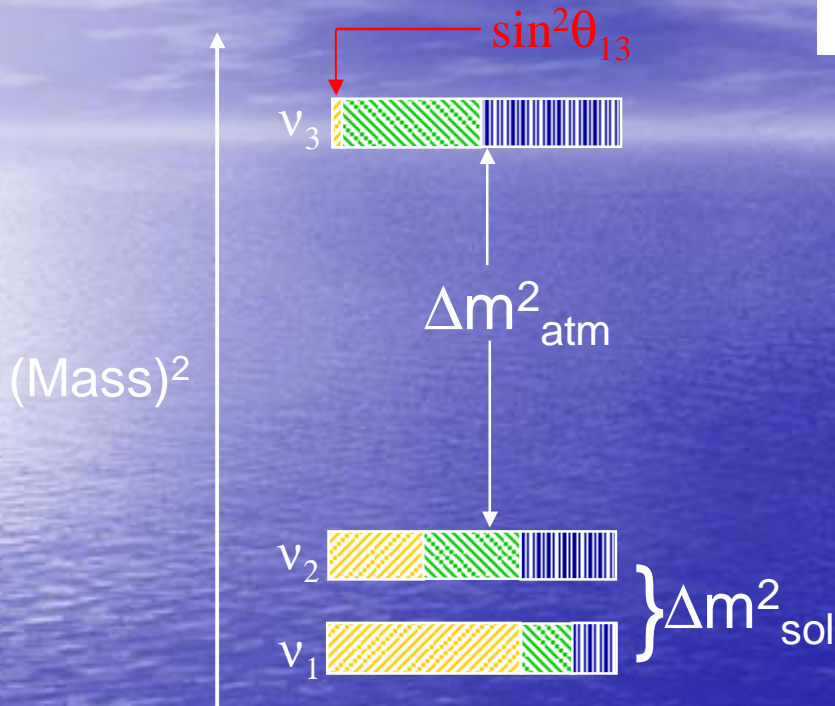
Tritium experiment

Onbb

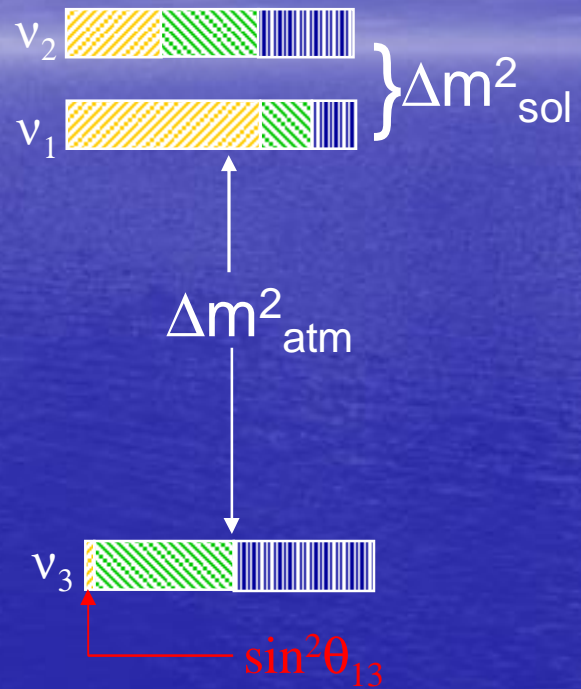
- value of  $\theta_{13}$
- CP violation in lepton sector
- mass spectrum: normal or inverted
- neutrino mass
- Majorana/Dirac nature

# Mass spectrum

$\sin^2\theta_{13} = |U_{e3}|^2 \rightarrow$  small  $\nu_e$  piece of  $\nu_3$   
 $\nu_3$  is at one end of  $\Delta m^2_{\text{atm}}$   
 We need an experiment with L/E sensitive to  $\Delta m^2_{\text{atm}}$   
 (L/E  $\sim$  500 km/GeV) involving  $\nu_e$



or



$\nu_e [ |U_{ei}|^2 ]$

$\nu_\mu [ |U_{\mu i}|^2 ]$

$\nu_\tau [ |U_{\tau i}|^2 ]$

# Test of discrete symmetries

$$P(n_a \rightarrow n_b) \neq P(\bar{n}_b \rightarrow \bar{n}_a)$$

CPT violation

$$P(n_a \rightarrow n_a) \neq P(\bar{n}_a \rightarrow \bar{n}_a)$$

$$P(n_a \rightarrow n_b) \neq P(\bar{n}_a \rightarrow \bar{n}_b)$$

CP violation

$$P(n_a \rightarrow n_b) \neq P(\bar{n}_b \rightarrow \bar{n}_a)$$

T violation

$$A_{CP} = \frac{P(n_m \textcircled{R} n_e) - P(\bar{n}_m \textcircled{R} \bar{n}_e)}{P(n_m \textcircled{R} n_e) + P(\bar{n}_m \textcircled{R} \bar{n}_e)} @ \frac{Dm_{12}^2 L}{4E_n} \times \frac{\sin 2q_{12}}{\sin q_{13}} \times \sin d$$

For  $q_{12}=p/8$   $Dm_{12}^2=7 \times 10^{-5}$   $\sin^2 q_{12}=0.01$  (1/10 of CHOOZE limit)  $q = p/4$

$A_{CP} = 25\%$



# LBL experiment T2K (Tokai to Kamioka)

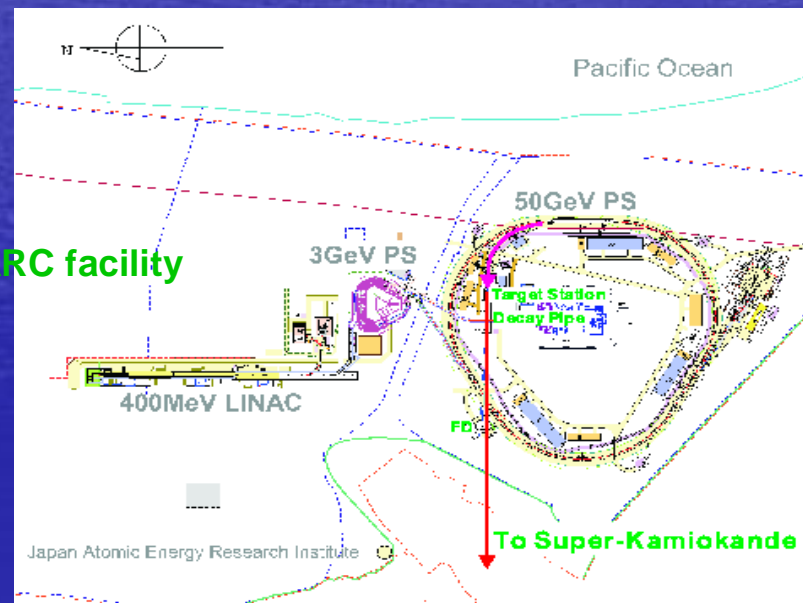
12 countries, 60 institutions, ~180 collaborators



$\nu$  beam    off-axis    on-axis

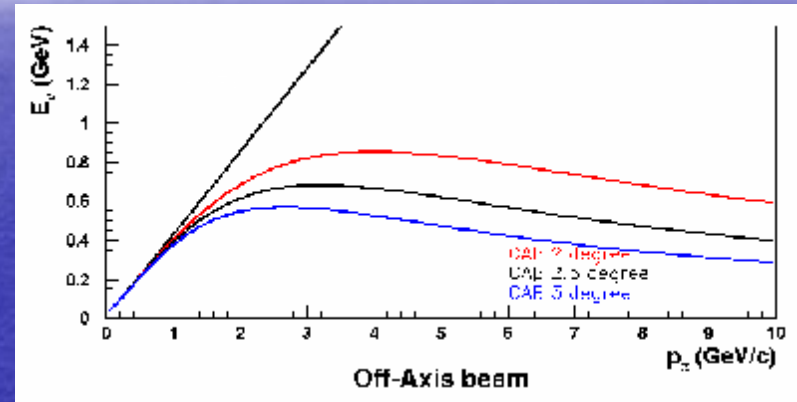
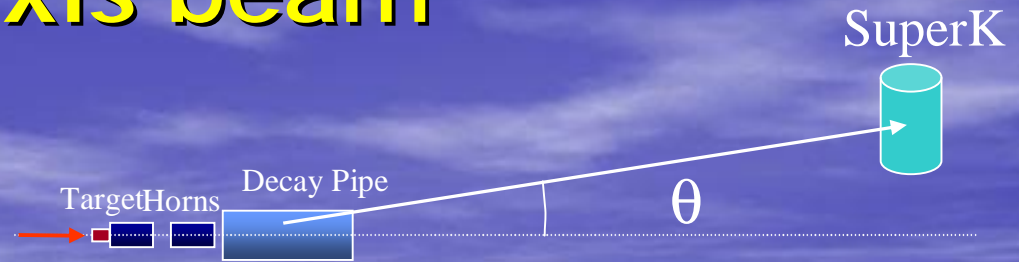
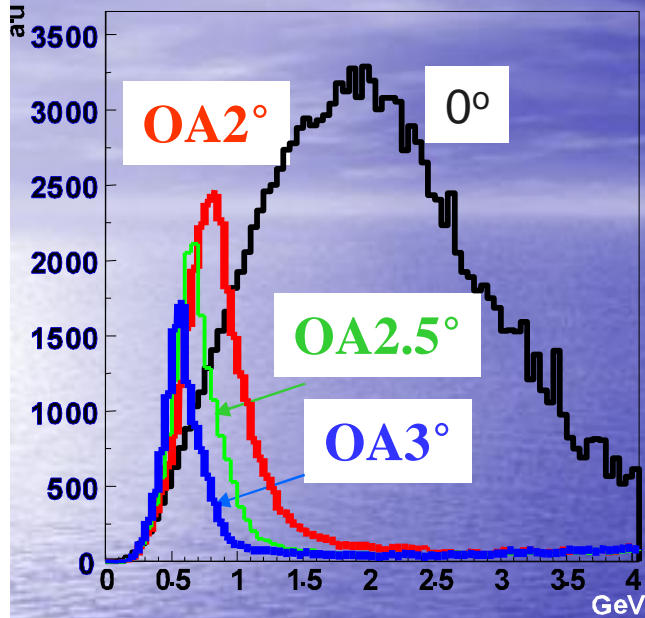
	JPARC	MINOS	K2K
E(GeV)	50	120	12
Int( $10^{12}$ ppp)	330	40	6
Rate (Hz)	0.29	0.53	0.45
Power (MW)	0.77	0.41	0.0052

JPARC facility



$\sim 1\text{GeV } n_m$  beam ( $\sim 100$  of K2K)

# Off-axis beam



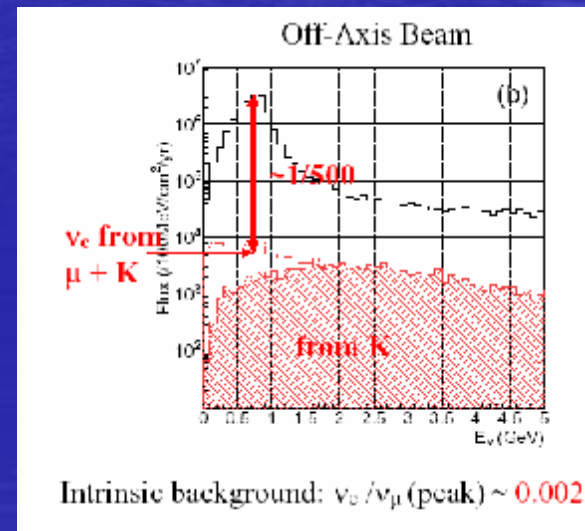
## Statistics at SK

OAB 2.5 deg, 1 yr =  $10^{21}$  POT, 22.5 kt

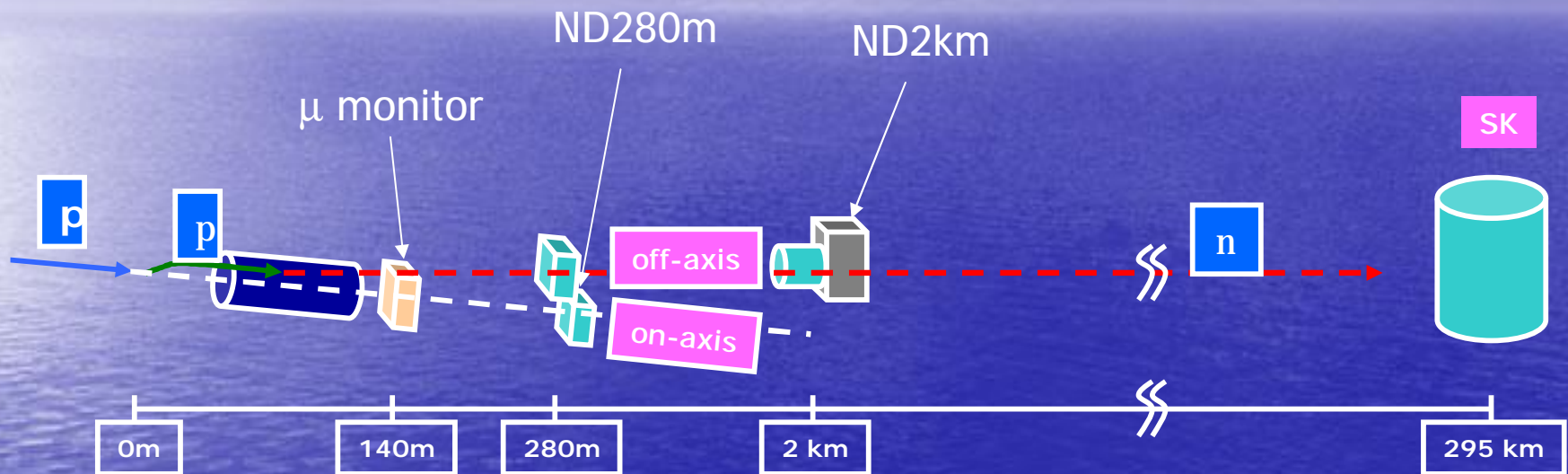
~ 2200  $n_m$  tot

~ 1600  $n_m$  charged current

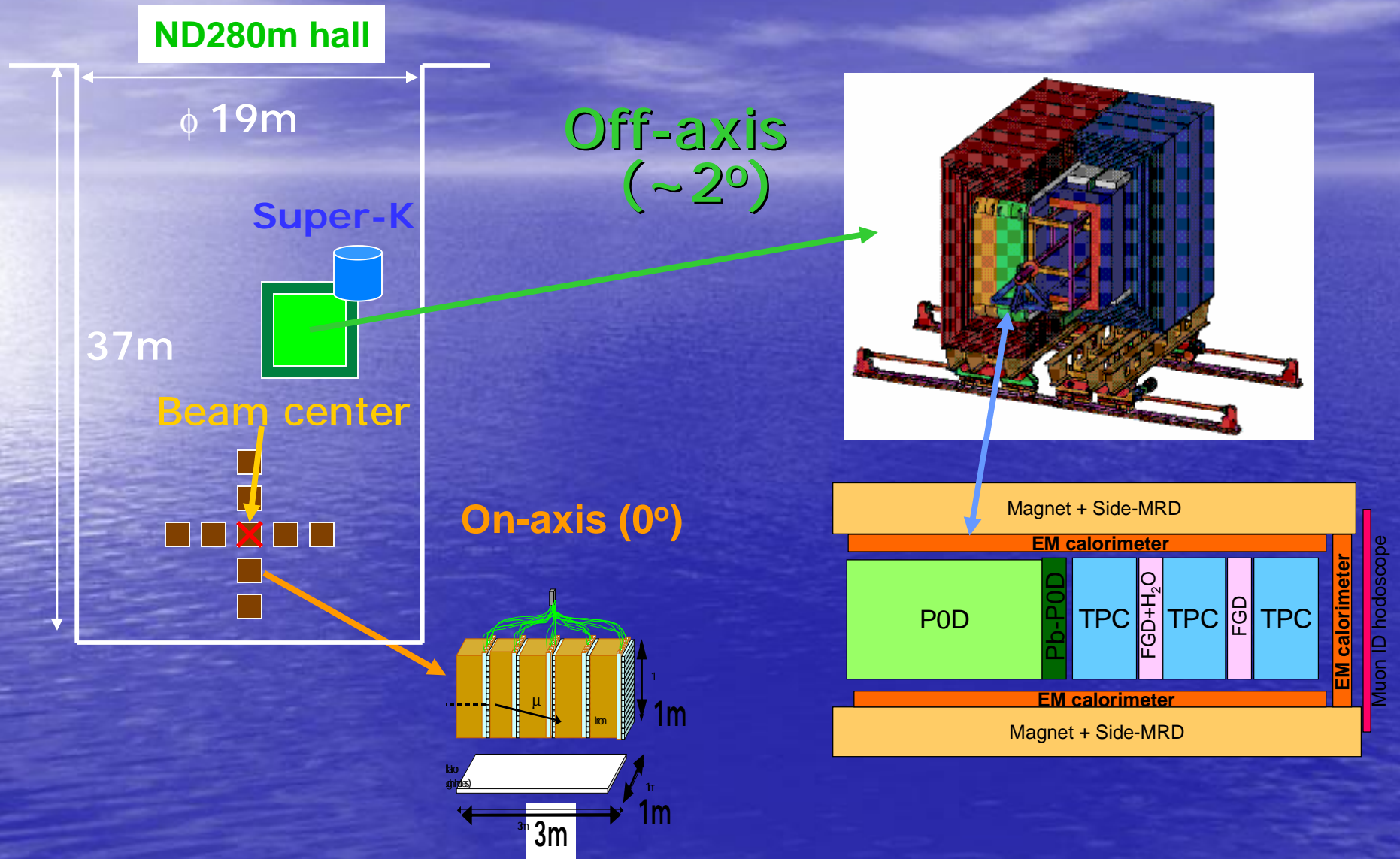
$n_e < 0.5\%$  at  $n_m$  peak



# T2K detectors

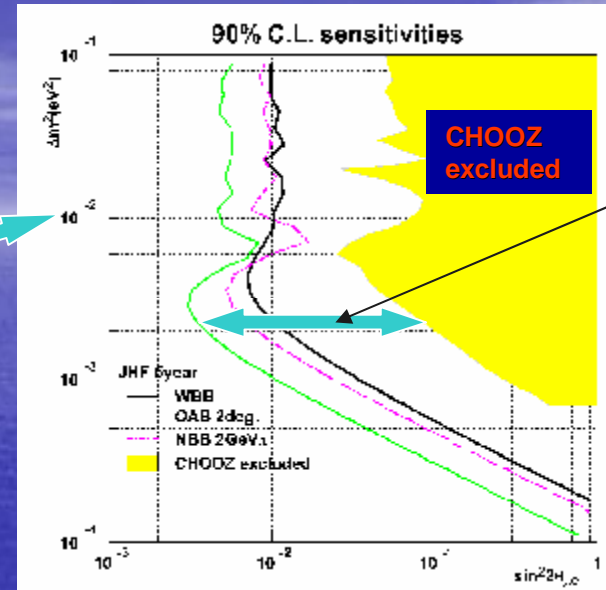
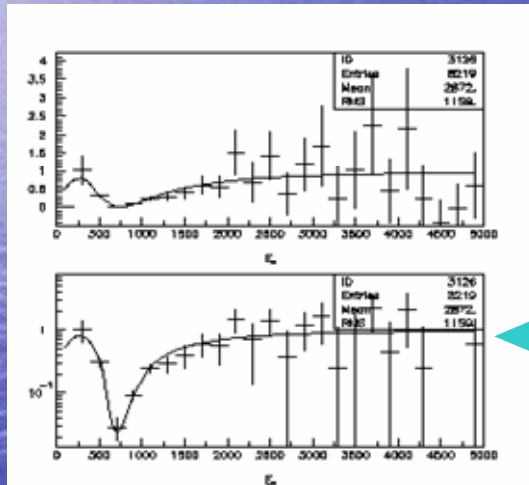


# Near Detectors at 280m



# Principle Goals of T2K

- Search for  $\nu_e$  appearance  
 $q_{13}$  sensitivity  $\pm 1^\circ$  (90% c.l.)



- Measurement  $Dm^2_{23}$  with accuracy of 1%  
 $d(\sin^2 2q_{23}) \sim 0.01$   
 $d(Dm^2_{23}) < 1 \cdot 10^{-4} \text{ eV}^2$

- Confirmation of  $\nu_m \leftrightarrow \nu_t$  oscillation

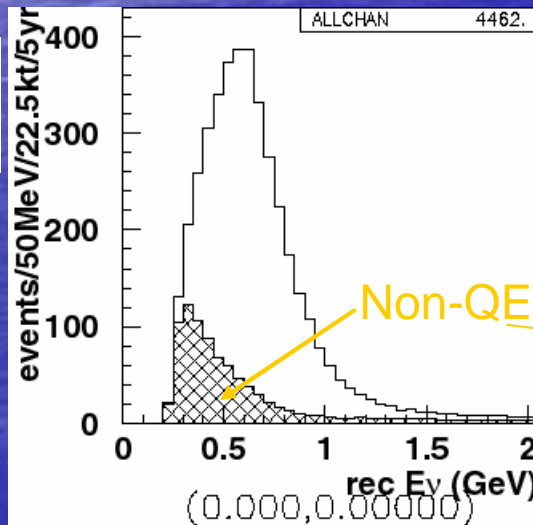
# Number of $n$ events at SK for 5 years

Off-axis (deg)	w/o oscillation (events/22.5kt/5yr)	max. deficit	$\Delta m^2$ (eV <sup>2</sup> )
2.0	6683	1724	$3.22 \times 10^{-3}$
2.5	4462	1103	$2.70 \times 10^{-3}$
3.0	3006	752	$2.33 \times 10^{-3}$

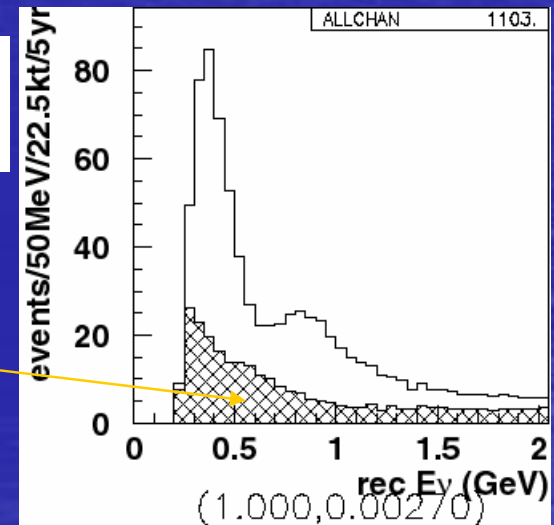
( $\sin^2 2\theta=1$ )

## Reconstructed $E_\nu$ spectrum at SK (OA2.5deg)

no  
oscillation



( $\sin^2 2\theta, \Delta m^2$ )  
= (1.0, 0.0027)



# LBL experiments

2006 -08

MINOS (FNAL)

$\nu_{\mu} \rightarrow \nu_{\mu}$  search for  $\nu_{\mu} \rightarrow \nu_e$

OPERA (CERN/Gran Sasso)

search for  $\nu_{\mu} \rightarrow \nu_{\tau}$  appearance

MiniBooNe(FNAL)

LSND anomaly

2009....

T2K Phase I

search for  $\nu_{\mu} \rightarrow \nu_e$  appearance/ $\theta_{13}$  measurement

Phase II **depends on  $\theta_{13}$**

CP – violation, if  $\theta_{13} \neq 0$

NOVA (FNAL)

CP – violation, mass hierarchy

# Summary

- Neutrinos have masses
- Clear signal of New Physics beyond the Standard Model (Solar, atmospheric, accelerator experiments)
- Exciting physics from running and future long base-line experiments
  - search/measurement of  $\theta_{13}$   $|U_{e3}| = ?$
  - precision measurements of  $\theta_{23}$  and  $\Delta m^2_{23}$
  - CP violation if  $\theta_{13}$  is large
  - mass hierarchy
- Unexpected or exotic properties?