

Particle and AstroParticle physics in China.

Status and perspectives. Part I

*Based on a talk of Yifang Wang
at 20th Lomonosov Conf., Aug. 19, 2021*

By Oleg Fedin



BEPCII - Beijing Electron Positron Collider II

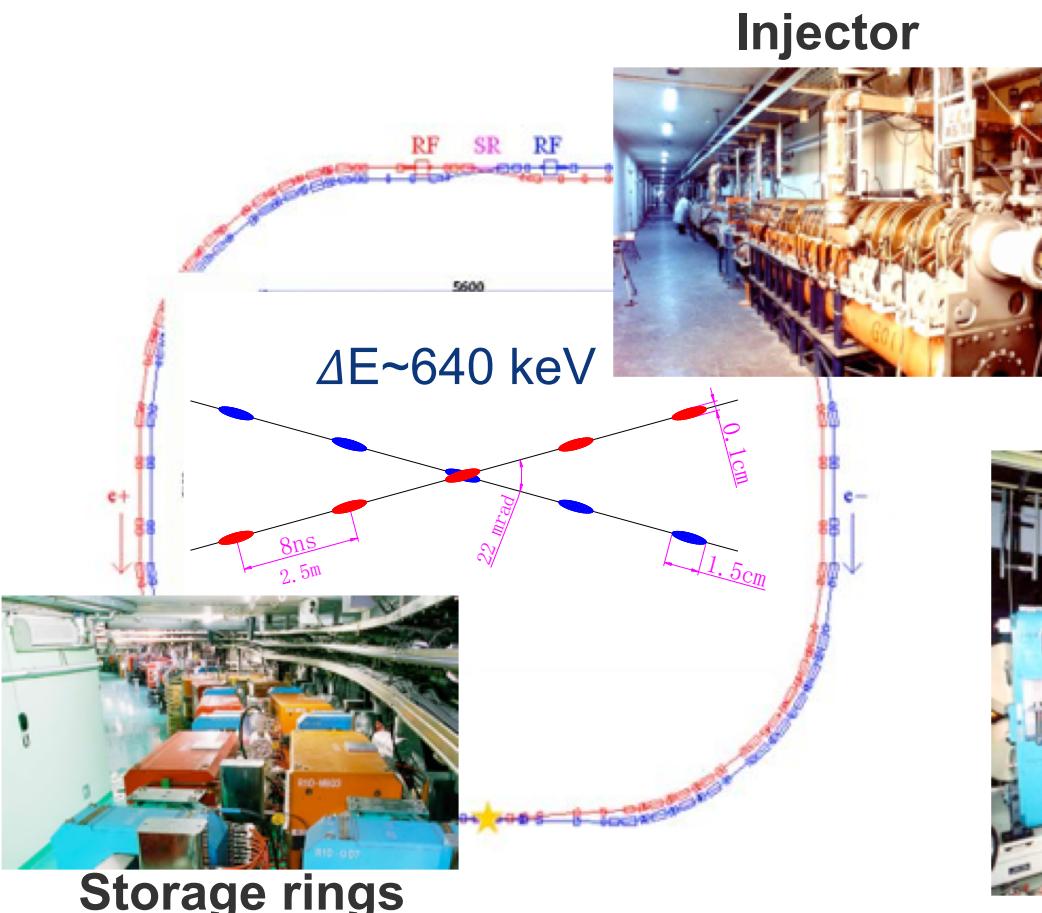
An e^+e^- collider at $2 - 4.6$ GeV

Design luminosity $10^{33} \text{ cm}^{-2}\text{s}^{-1}$ @1.89 GeV

Beam lifetime 2.7 hrs.

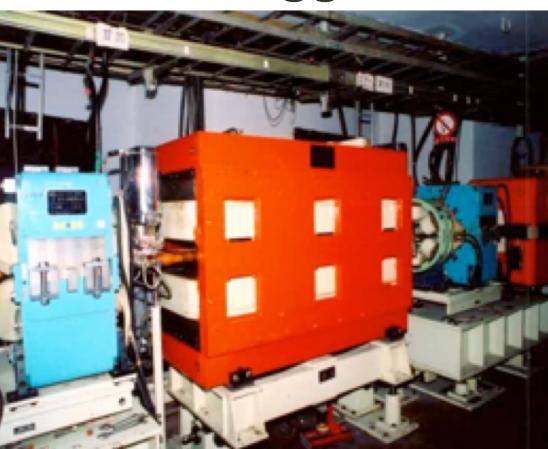
Operational from 2009

Highest peak and integral lumi recorded in 2020



It is comparable with all the second-generation synchrotron radiation light sources running in the world.

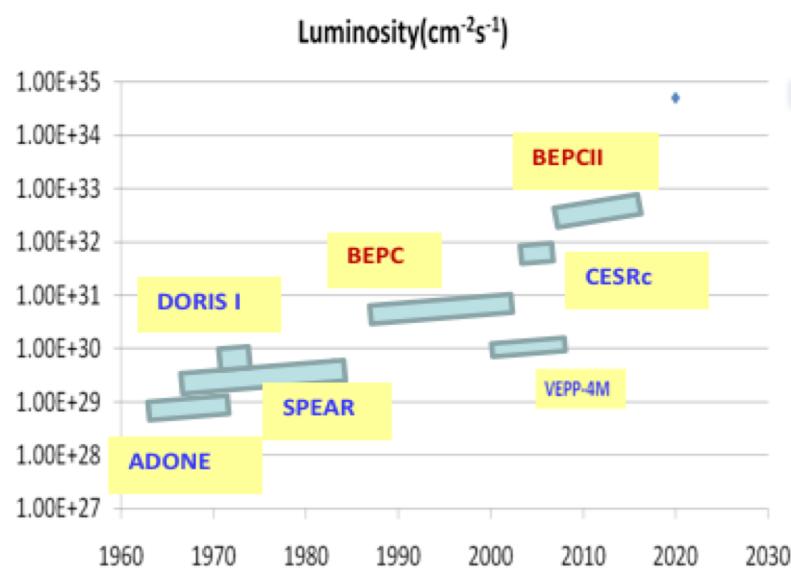
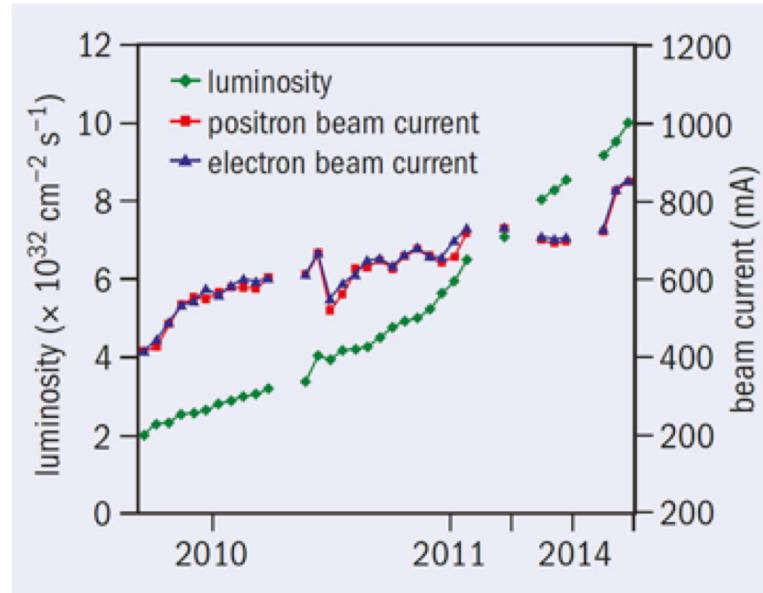
wiggler



SR front end region



BEPCII - Beijing Electron Positron Collider II



arXiv:0911.4960

Expected data samples in a calendar year based on assumption that the total running time will be 10^7 sec/year . Accordingly integrated lumi per year $\sim 1 \text{ fb}^{-1}$

States	Energy (GeV)	Peak luminosity ($10^{33} \text{ cm}^{-2} \text{ s}^{-1}$)	Physics cross-section (nb)	Events/year
J/ψ	3.097	0.6	3,400	1×10^{10}
$\psi(2S)$	3.686	1.0	640	3×10^9
$\tau^+ \tau^-$	3.670	1.0	2.4	1.2×10^7
$D^0 \bar{D}^0$	3.770	1.0	3.6	1.8×10^6
$D^+ D^-$	3.770	1.0	2.8	1.4×10^6
$D_s D_s$	4.030	0.6	0.32	1×10^6
$D_s \bar{D}_s$	4.170	0.6	1.0	2×10^6

σ -factory @ Novosibirsk

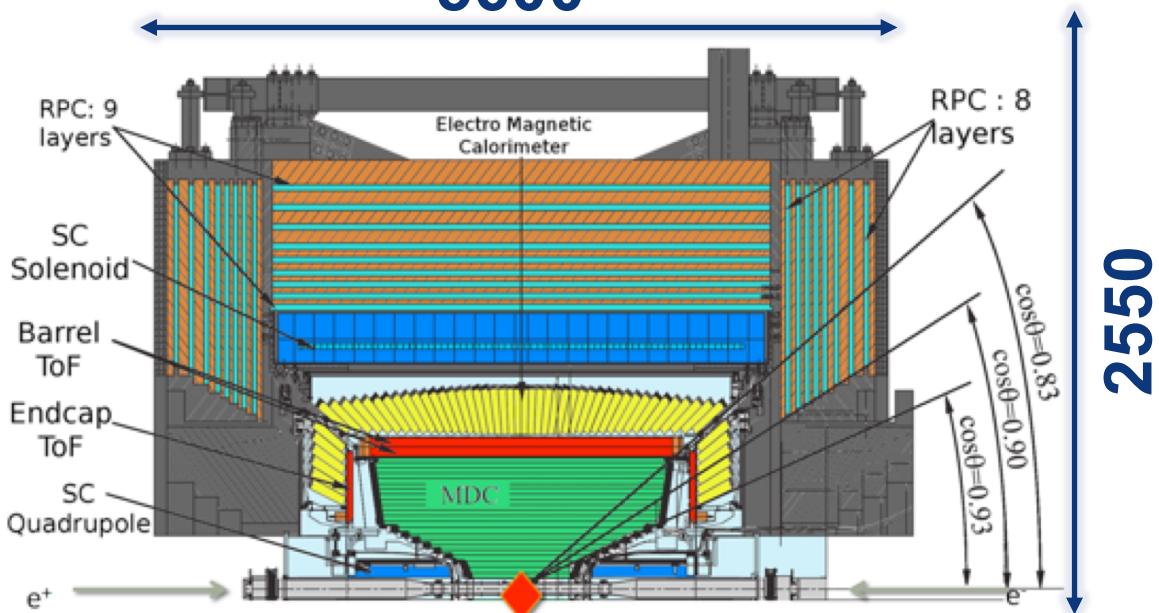
BES III (*Beijing Spectrometer III*)

5600

02/2019: 10^{10} J/ψ
Collected around 30 fb^{-1}

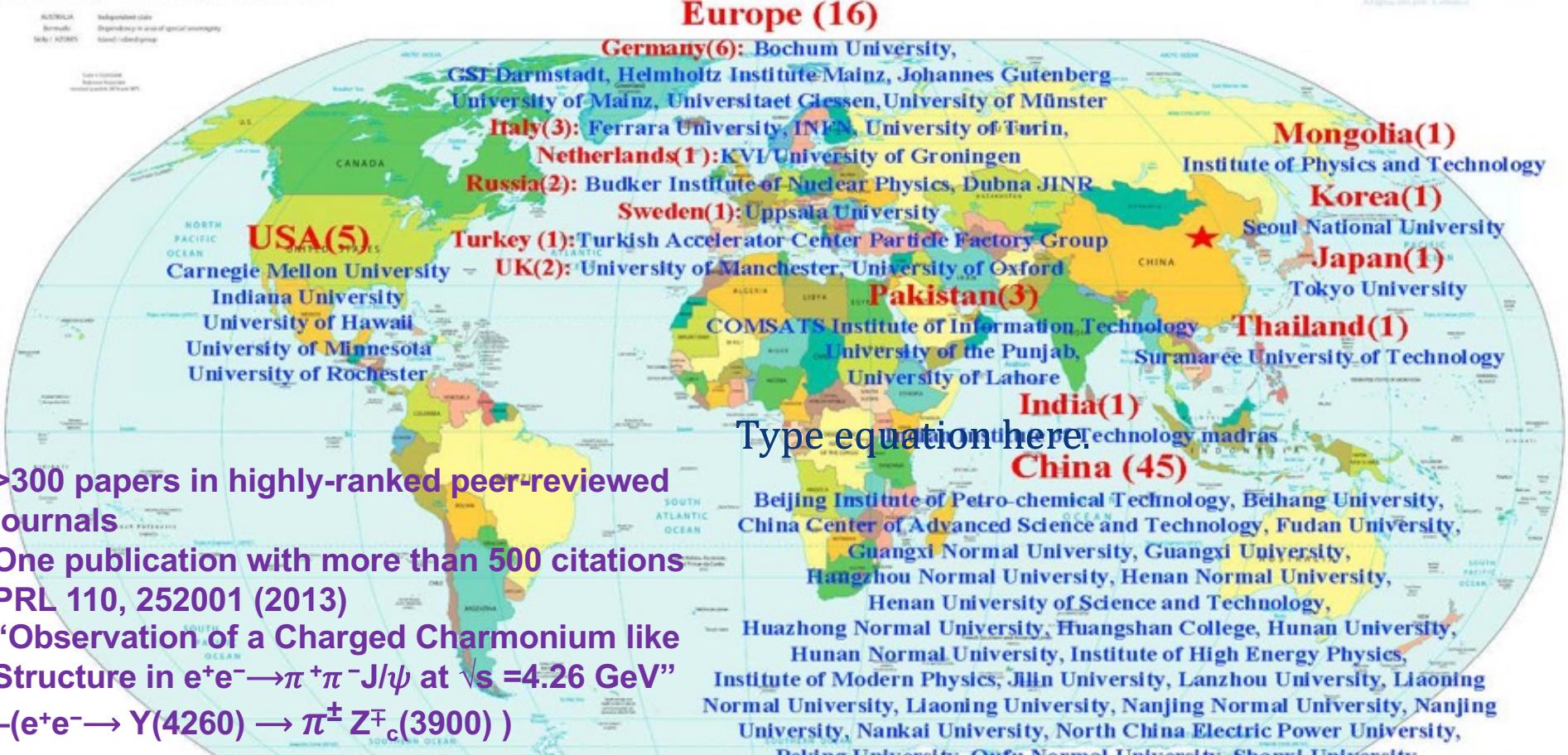
arXiv:0911.4960

Sub-system	BESIII
Single wire σ_{ϕ} (μm)	130
MDC	σ_p/p (1 GeV/c)
	0.5%
	σ (dE/dx)
	6 %
EMC	σ_E/E (1GeV)
	2.5%
	Position resolution (1 GeV)
	0.6 cm
TOF	σ_T (ps)
	Barrel
	100
	End cap
	110
Muon	No. of layers (barrel/end cap)
	9/8
	cut-off momentum (MeV/c)
	0.4
	Solenoid magnet Field (T)
	1.0
	$\Delta\Omega/4\pi$
	93%



The BES III Collaboration

Political Map of the World, November 2011



>300 papers in highly-ranked peer-reviewed journals

One publication with more than 500 citations
PRL 110, 252001 (2013)

“Observation of a Charged Charmonium like Structure in $e^+e^- \rightarrow \pi^+\pi^-J/\psi$ at $\sqrt{s} = 4.26$ GeV”
 $-(e^+e^- \rightarrow Y(4260) \rightarrow \pi^\pm Z_c^\mp(3900))$



~500 members

From 74 institutions in 15 countries

$Y(4260)$ discovery: 2005 BaBar $B^- \rightarrow Y(4260)K^-$

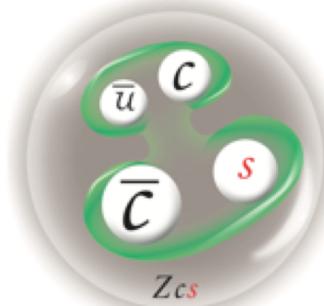
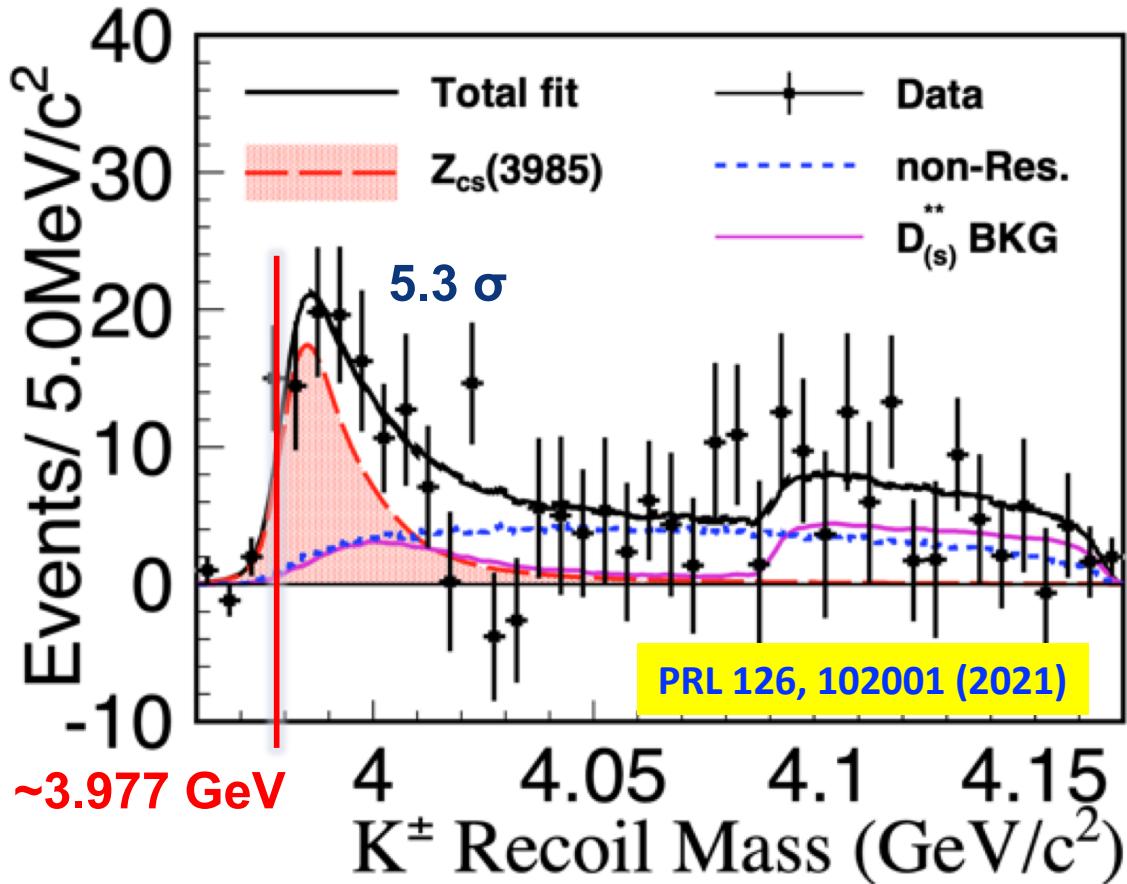


12/10/21

Семинар ОФВЭ

BES III: Observation of strange four-quark meson $Z_{cs}(3985)$

$e^+ e^- \rightarrow K^+ (D_s^- D^{*0} + D_s^{*-} D^0)$ @4.68 GeV



$e^+ e^- \rightarrow K^+ Z_{cs}(3985)$

$M = 3982.5^{+1.8}_{-2.6} \pm 2.1 \text{ MeV}$

$\Gamma = 12.8^{+5.3}_{-4.4} \pm 3.0 \text{ MeV}$

$J^P = 1^-$

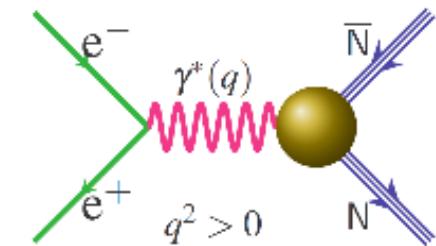
$D_s^- \rightarrow K^+ K^- \pi^-$ & $D_s^- \rightarrow K^0_S K^-$

BES III: Damped oscillatory behavior in neutron form-factor

$$\sigma_B = \frac{4\pi\alpha_{em}^2\beta C(q^2)}{3q^2} \left[|G_M(q^2)|^2 + |G_E(q^2)|^2 \frac{1}{2\tau} \right]$$

$$|G| = \sqrt{\frac{2\tau|G_M(q^2)|^2 + |G_E(q^2)|^2}{2\tau + 1}}$$

$$\tau = q^2/4m_N^2$$



Phys. Rev. Lett. 114, 232301 (2015)

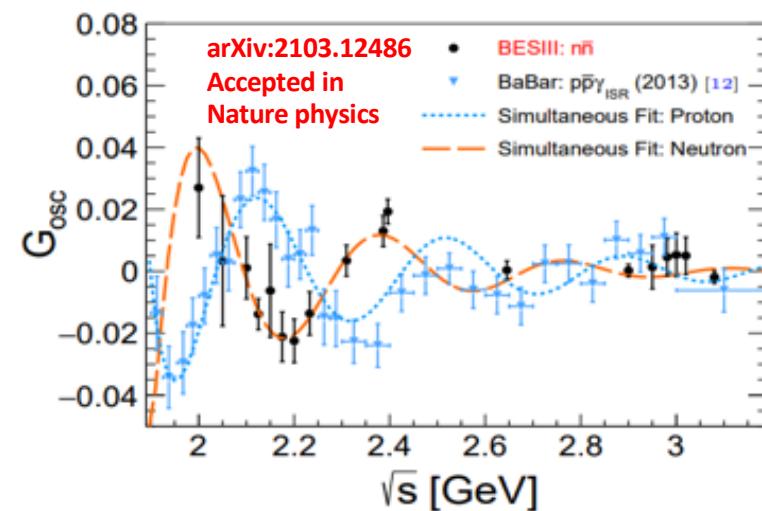
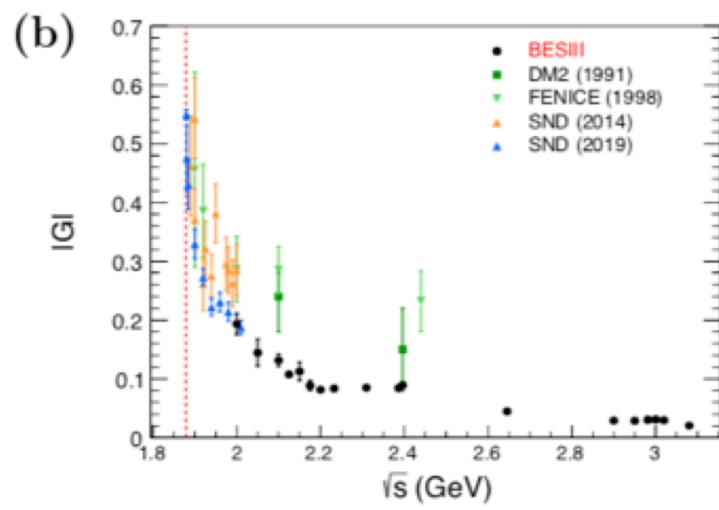
$$G(q^2) = G_0(q^2) + G(q^2)_{osc}$$

regular background fit oscillation fit

$$|G_0^p(q^2)| = \frac{A_p}{(1 + q^2/m_a^2)(1 - q^2/0.71)}$$

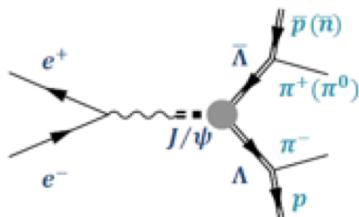
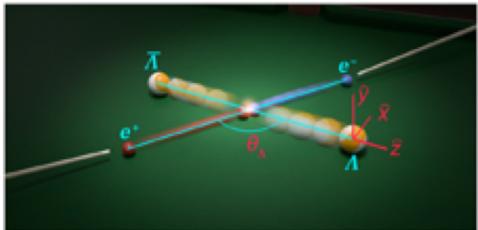
$$|G_p^n(q^2)| = \frac{A_n}{(1 - q^2/0.71)^2}$$

$$G(p)_{osc}^{n,p} = A^{n,p} e^{-B^{n,p} p} \cos(Cp + D^{n,p}) \quad p \equiv \sqrt{E^2 - m_{n,p}^2} \quad E = \frac{q^2}{2m_{n,p}} - m_{n,p}$$



It suggests an interference effect involving rescattering processes at moderate kinetic energies of the outgoing hadrons. Such processes take place when the centres of mass of the produced hadrons are separated by ≈ 1 fm.

BES III: *Polarization and entanglement in baryon–antibaryon pair production e^+e^- annihilation*

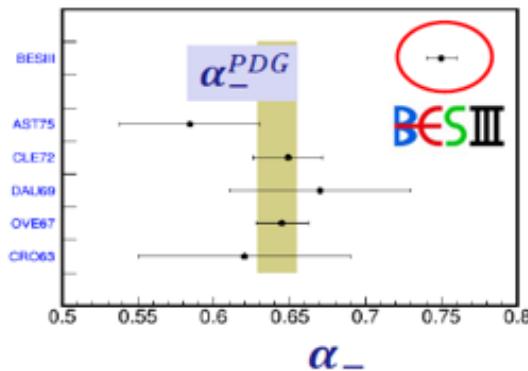


Parameters	This work	Previous results
α_ψ	$0.461 \pm 0.006 \pm 0.007$	0.469 ± 0.027 (ref. ¹⁴)
$\Delta\Phi$	$42.4 \pm 0.6 \pm 0.5^\circ$	-
α_-	$0.750 \pm 0.009 \pm 0.004$	0.642 ± 0.013 (ref. ⁶)
α_+	$-0.758 \pm 0.010 \pm 0.007$	-0.71 ± 0.08 (ref. ⁶)
$\bar{\alpha}_0$	$-0.692 \pm 0.016 \pm 0.006$	-
A_{CP}	$-0.006 \pm 0.012 \pm 0.007$	0.006 ± 0.021 (ref. ⁶)
$\bar{\alpha}_0/\alpha_+$	$0.913 \pm 0.028 \pm 0.012$	-

$\Lambda \rightarrow p\pi^-$: $\alpha_- = 0.750 \pm 0.009 \pm 0.004$

Hyperon up-down parameter: 7σ upward shift from all previous measurements

Nature Physics 15, 631–634(2019)



CP asymmetry SM predict $\sim 10^{-4}$

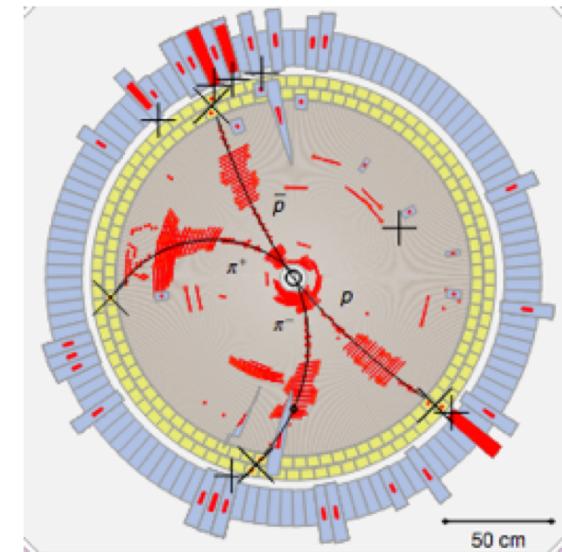
$$A_{CP} = (\alpha_- + \alpha_+)/(\alpha_- - \alpha_+)$$

$$\frac{dN}{d\Omega} = \frac{1}{4\pi} (1 + \alpha \vec{P}_\Lambda \vec{n}), \quad |\alpha| < 1$$

$$J/\psi \rightarrow (\Lambda \rightarrow p\pi^-)(\bar{\Lambda} \rightarrow \bar{p}\pi^+)$$

$$\mathcal{W}(\xi; \alpha_\psi, \Delta\Phi, \alpha_-, \alpha_+)$$

$$\begin{aligned} &= 1 + \alpha_\psi \cos^2 \theta_\Lambda + \alpha_- \alpha_+ [\sin^2 \theta_\Lambda (n_{1,x} n_{2,x} - \alpha_\psi n_{1,y} n_{2,y}) \\ &+ (\cos^2 \theta_\Lambda + \alpha_\psi) n_{1,z} n_{2,z}] \\ &+ \alpha_- \alpha_+ \sqrt{1 - \alpha_\psi^2} \cos(\Delta\Phi) \sin \theta_\Lambda \cos \theta_\Lambda (n_{1,x} n_{2,z} + n_{1,z} n_{2,x}) \\ &+ \sqrt{1 - \alpha_\psi^2} \sin(\Delta\Phi) \sin \theta_\Lambda \cos \theta_\Lambda (\alpha_- n_{1,y} + \alpha_+ n_{2,y}) \end{aligned}$$



$J/\psi \rightarrow \Lambda\bar{\Lambda}$: 400,000 events



BES III: Future Physics Program

□ Rich physics program requiring $> 40 \text{ fb}^{-1}$, corresponding to 15 yrs of data taking at the present luminosity

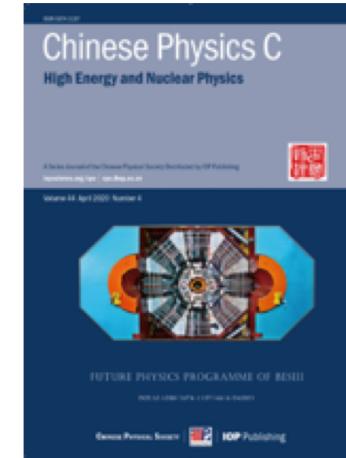
□ Detector upgrade:

- end-cap TOF (110 ps) => multi-gap resistive plate chamber (MRPC) 68 ps
- Inner MDC with a cylindrical GEM inner tracker

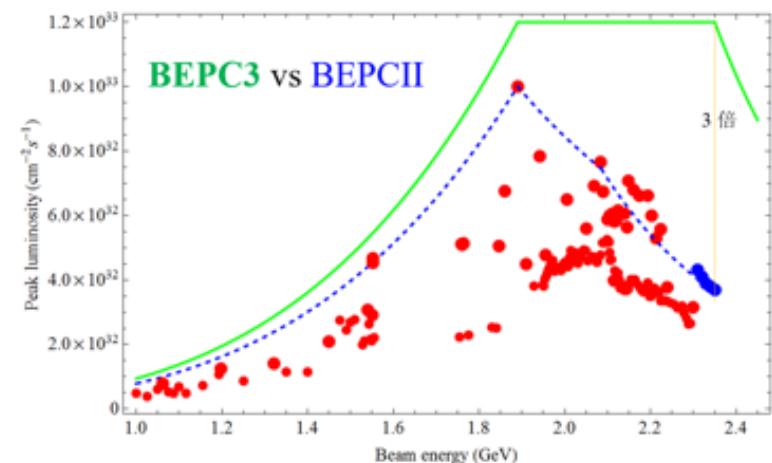
□ Collider upgrade:

- Luminosity $\times \sim 3 \rightarrow$ squeeze the beam size by adding a new RF cavity per beam
- Replace the two SC quadrupole magnets near the IP to increase the beam energy from 2.45 up to 2.8 GeV for the interests of studying Ξ_c
- slight decreasing energy down from 1 to 0.9 GeV, for the interests of studying nucleon productions

□ Operation will continue until at least 2030



Chinese Physics C 44, 040001 (2020)
arXiv:1912.05983v3

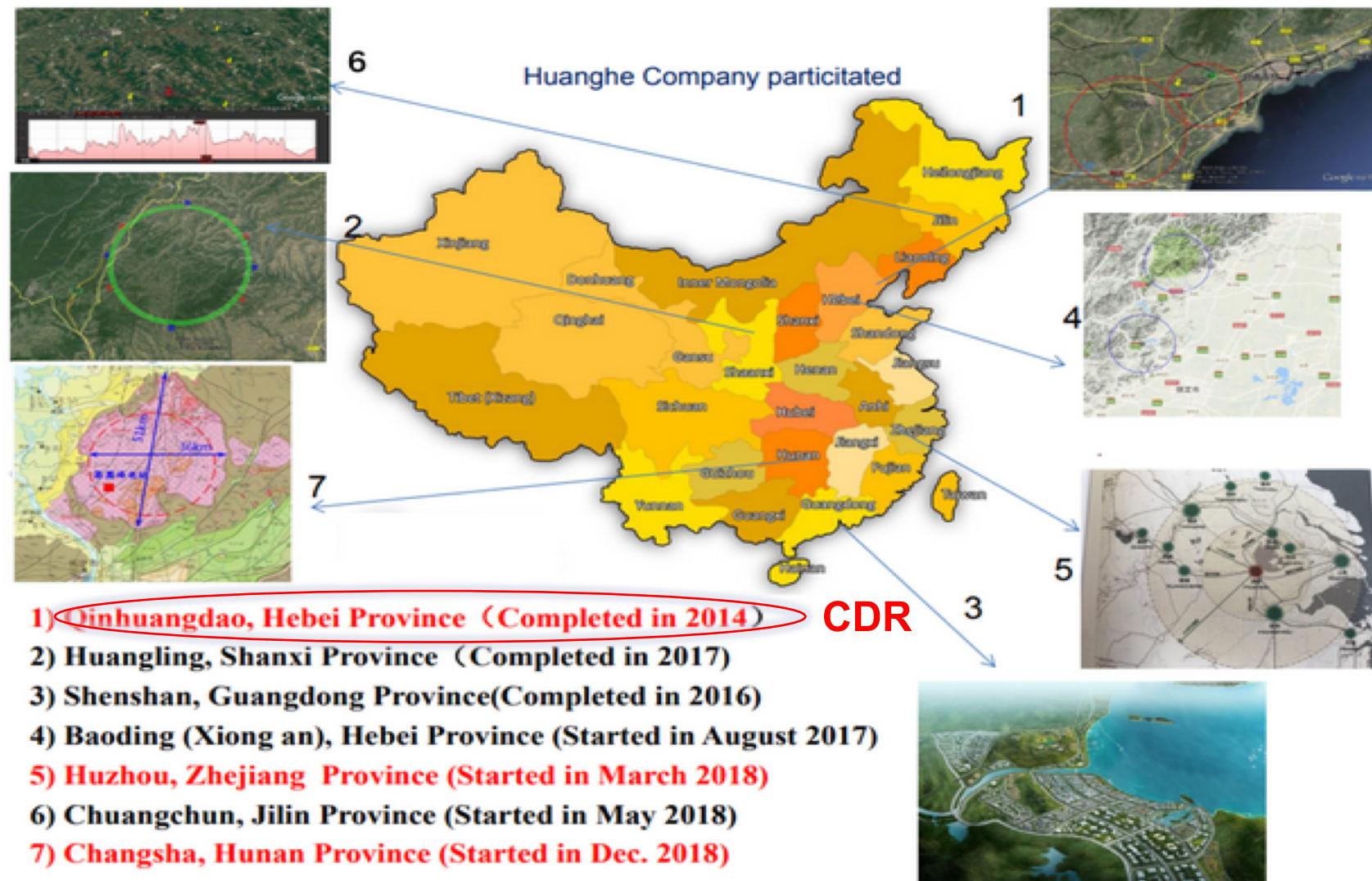


CEPC - Circular Electron Positron Collider

- Since 2005, we were discussing the next machine after BEPCII
- The idea of a Circular e+e- Collider (CEPC) followed by a Super proton-proton collider (SPPC) was proposed in Sep. 2012, and quickly gained the momentum in IHEP and in the world
 - Looking for Hints@ e^+e^- Collider → If yes, direct searches@PP collider
 - The tunnel can be re-used for pp, AA, ep colliders up to ~ 100 TeV → compatibility study needed now



CEPC Site selection

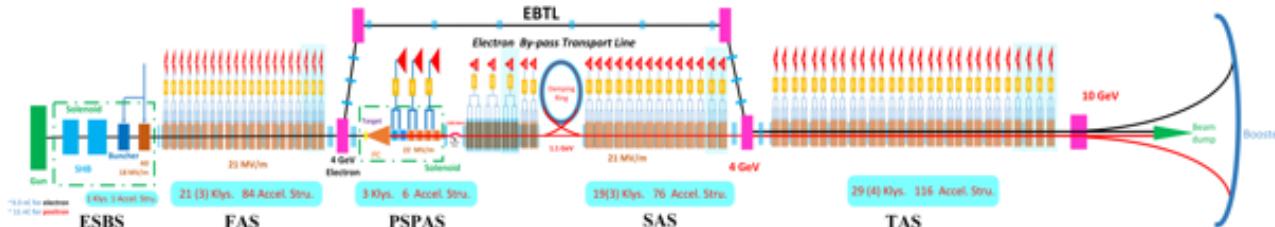


CEPC Possible timeline

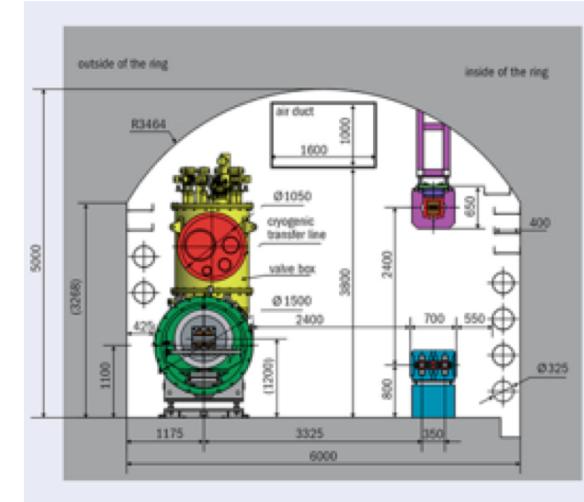


- The first stage is to complete a Preliminary Conceptual Design Report (PreCDR) in 2015 and a Conceptual Design Report (CDR) in 2018 (arXiv: 1809.00285 & arXiv:1811.10545). With the publication of this report, these goals have been achieved.
- The next stage is a 5-year period from 2018 to 2022 for R&D and for completion of a Technical Design Report (TDR)
- Construction will start in 2022 in the government's 14th Five-Year Plan and continue in the 15th Five-Year Plan. Construction will be completed by 2030.
- Experiments can begin as early as 2030 when the 16th Five-Year Plan starts.
- The experiments will continue for about 10 years until 2040.
- After 2040, the superconducting magnets for the SPPC project are expected to be ready for installation, and the SPPC era will begin.

CEPC Layout

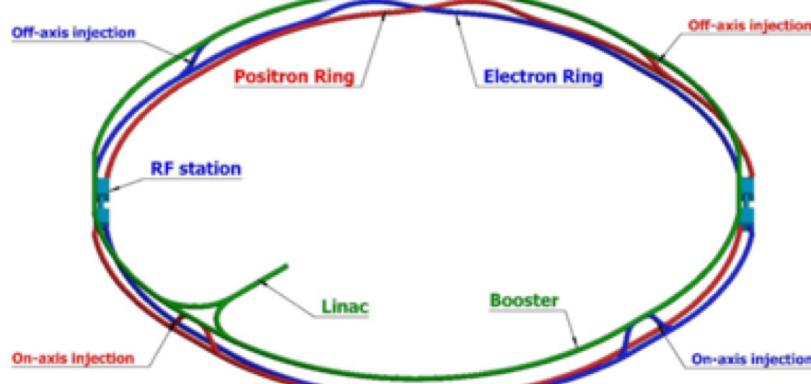


Linac injector (1.2km, 10GeV)

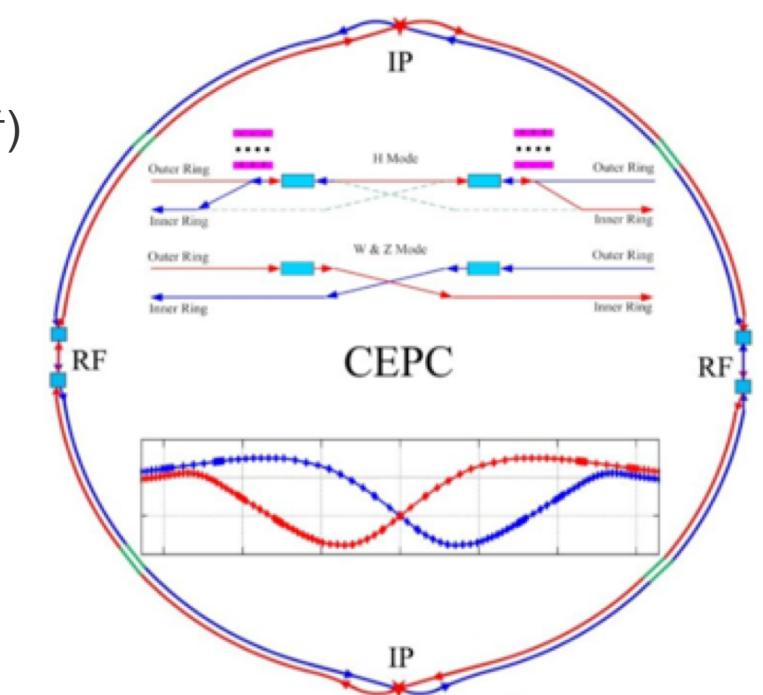


Collider ring (100km)

Booster ring (100km)



H ($e^+e^- \rightarrow ZH$)
Z ($e^+e^- \rightarrow Z$)
W ($e^+e^- \rightarrow W^+W^-$)



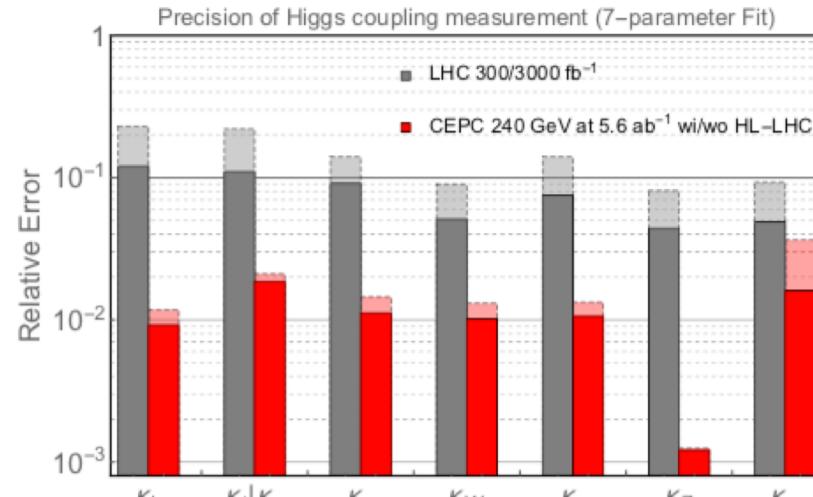
CEPC 10 years operation plan 7-2-1

1 атто = 10^{-18}

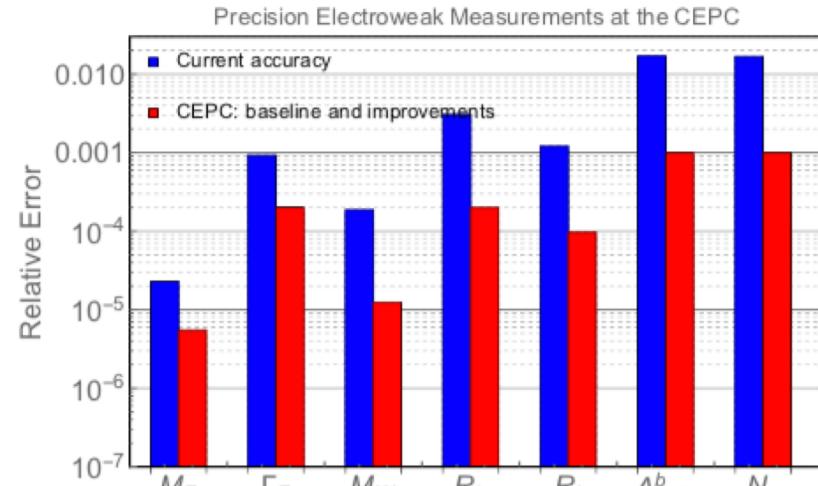
arXiv: 1809.00285

Particle	$E_{\text{c.m.}}$ (GeV)	L per IP ($10^{34} \text{ cm}^{-2}\text{s}^{-1}$)	Integrated L per year (ab^{-1} , 2 IPs)	Years	Total Integrated L (ab^{-1} , 2 IPs)	Total no. of particles
H	240	3	0.8	7	5.6	1×10^6
Z	91	32 (*)	8	2	16	7×10^{11}
W^+W^-	160	10	2.6	1	2.6	1.5×10^7

(*) Assuming detector solenoid field of 2 Tesla during Z operation



(a)



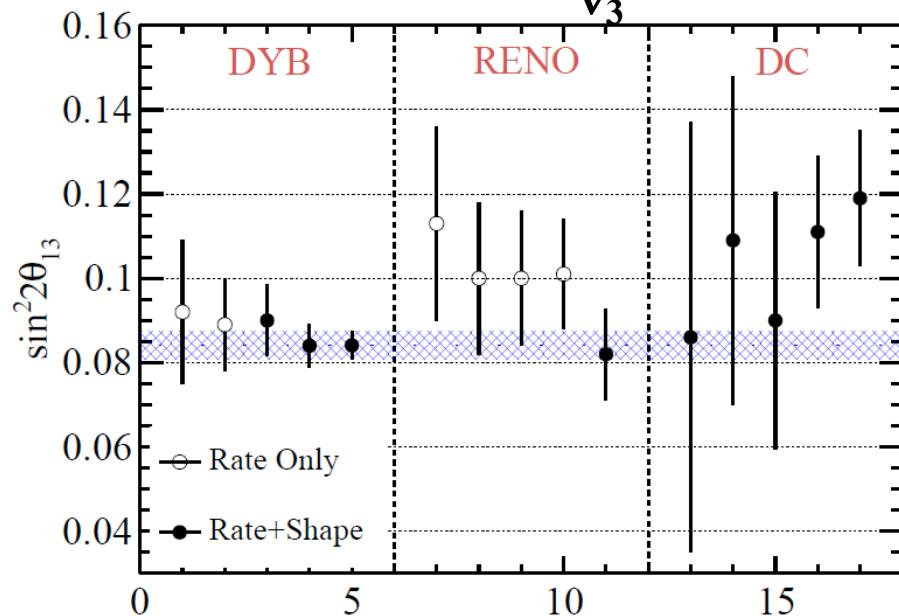
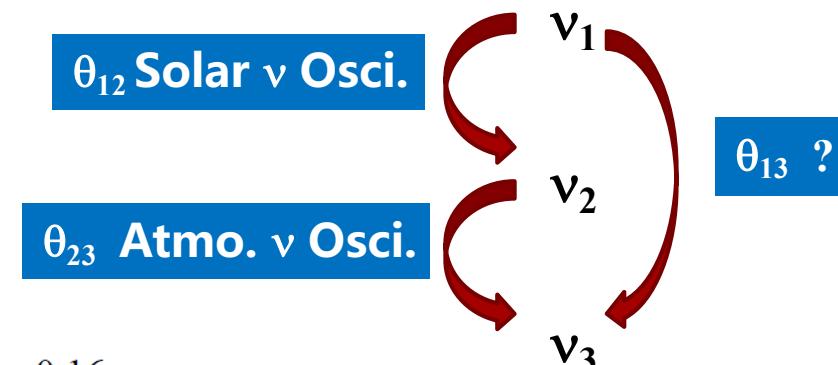
(b)

arXiv:1811.10545



Daya Bay Experiment

$$P = 1 - \sin^2(2\theta_{13}) \sin^2\left(\frac{1.27\Delta m_{31}^2 L}{E_\nu}\right)$$

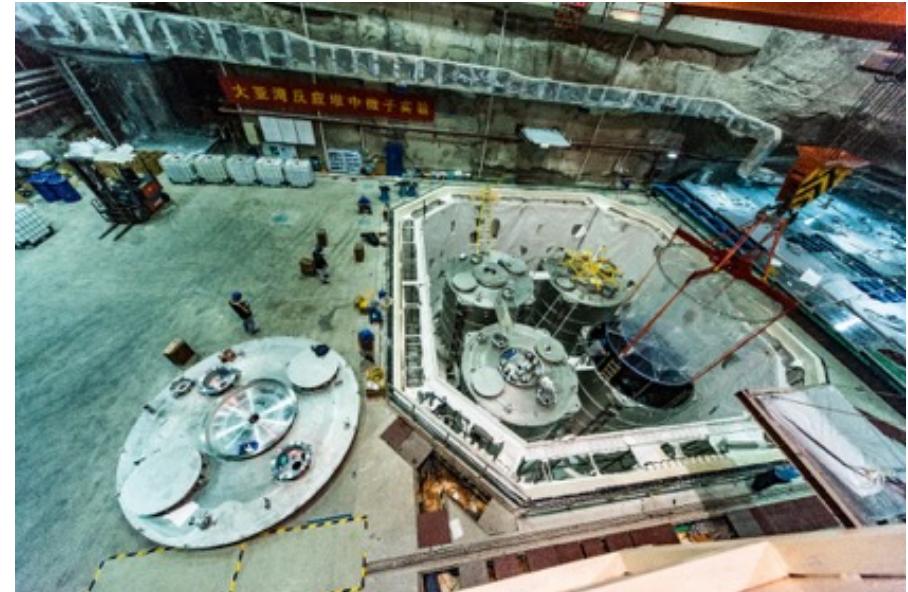


2012: $\sin^2(2\theta_{13}) = 0,092 \pm 0,016$ (стат.) $\pm 0,005$ (систем.)



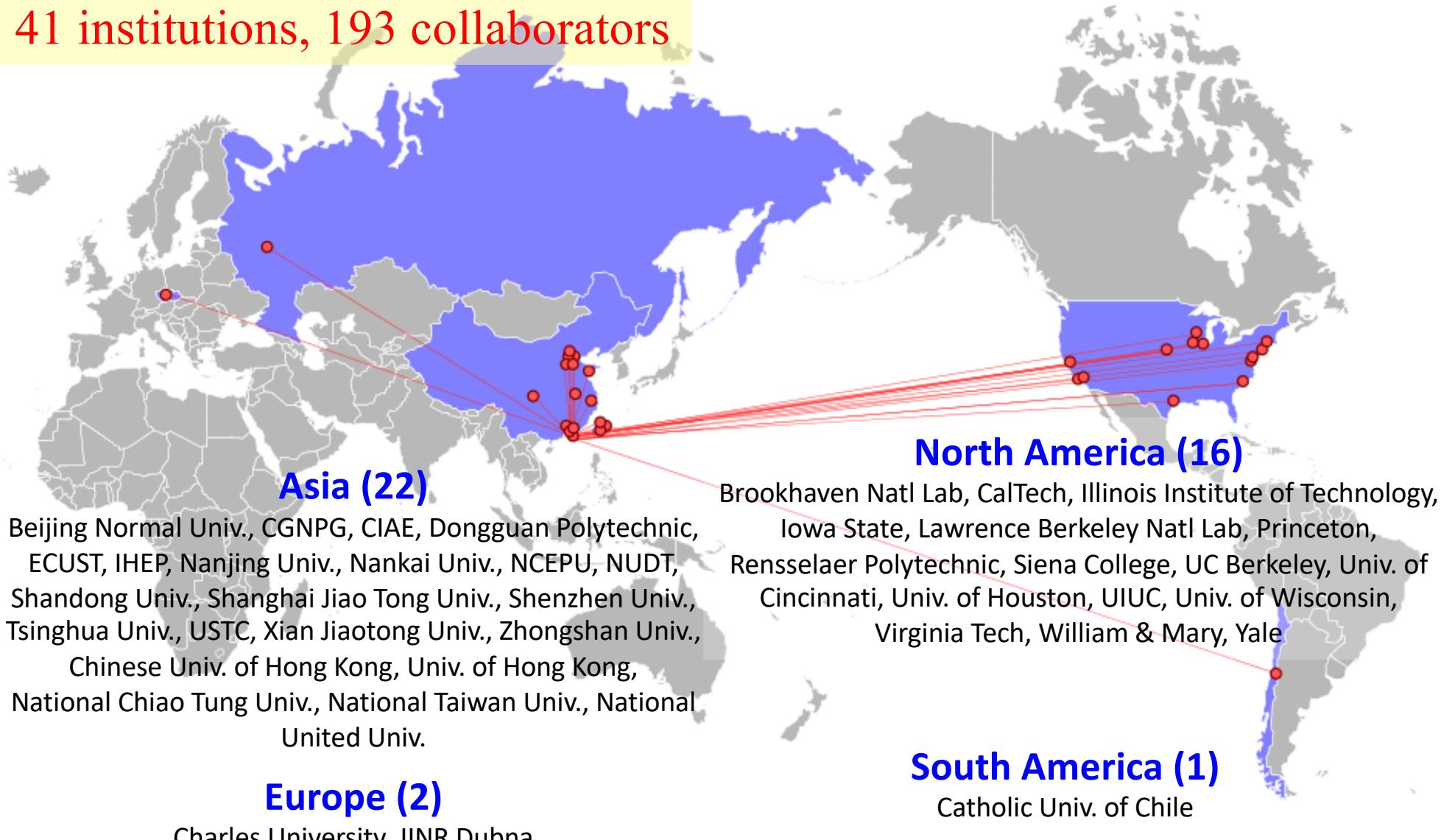
Daya Bay decommissioning

- Data taking terminated on Dec. 12, 2020
 - Ceremony broadcasted online, 1.7 M viewers
- Data analysis will continue, expected $\sin^2 2\theta_{13}$ precision < 3%
- Detector decommissioning in process



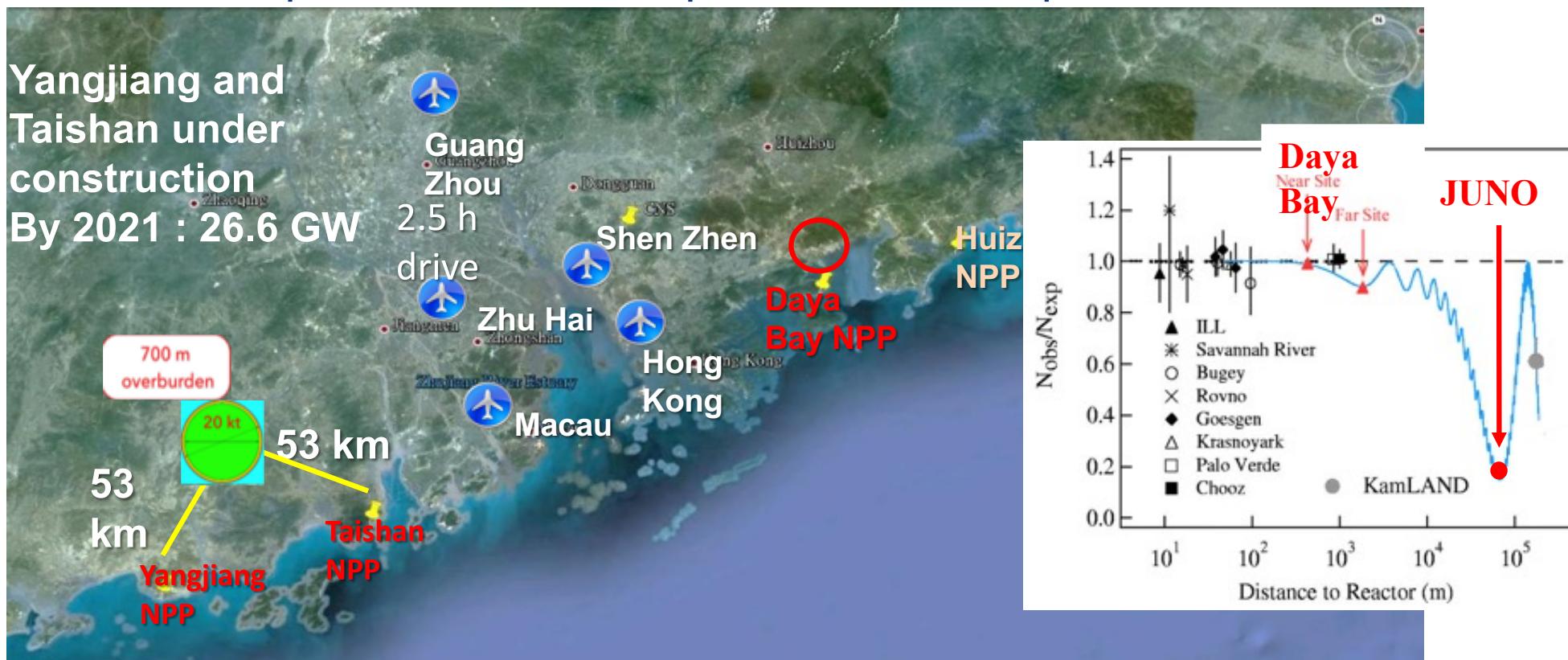
Daya Bay Collaboration

41 institutions, 193 collaborators



JUNO (*Jiangmen Underground Neutrino Observatory*) Experiment

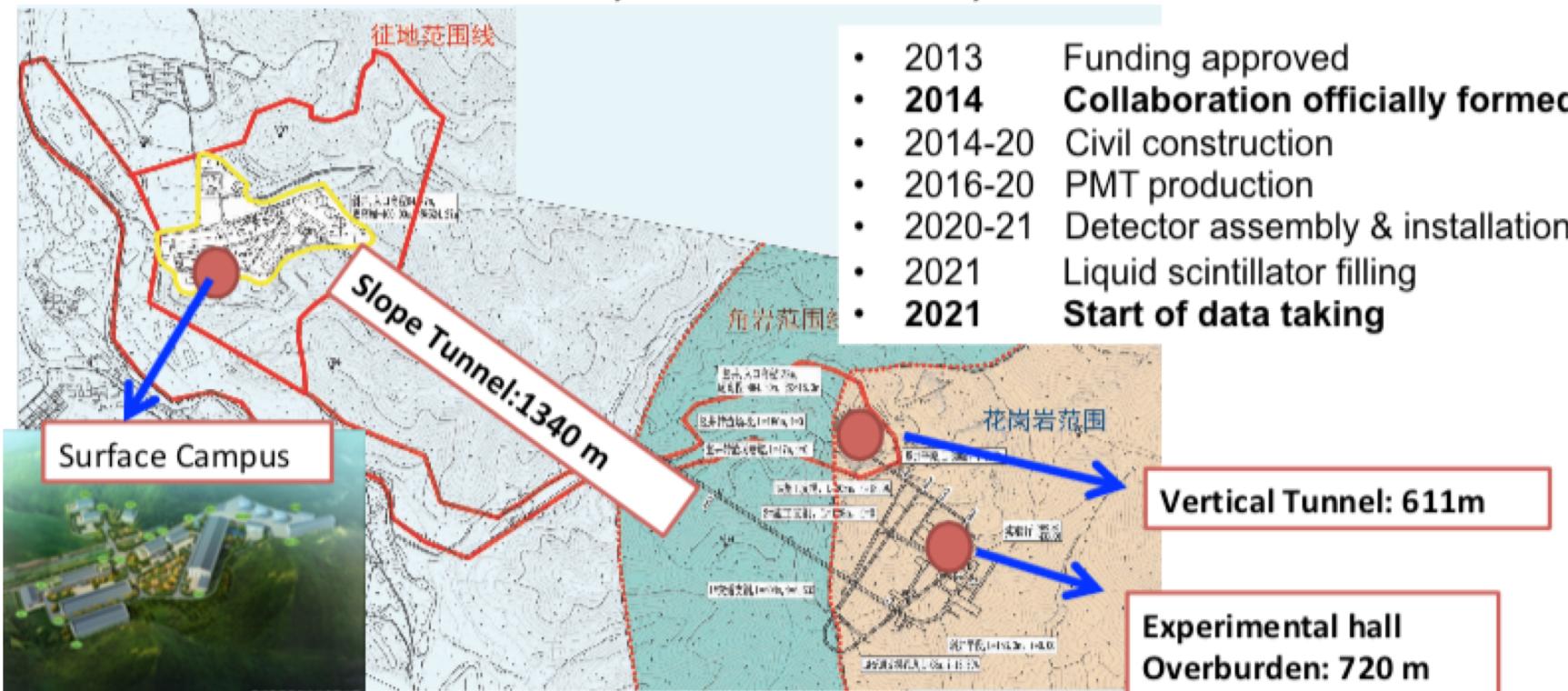
- JUNO is a “medium-baseline” (53km) reactor neutrino experiment (data taking foreseen in 2021)
- JUNO will be the largest Liquid Scintillator detector ever built (**20kt**)
- Goals : Measurement of the neutrino mass hierarchy (PNMH) and oscillation parameters + astroparticle and rare processes



JUNO Collaboration

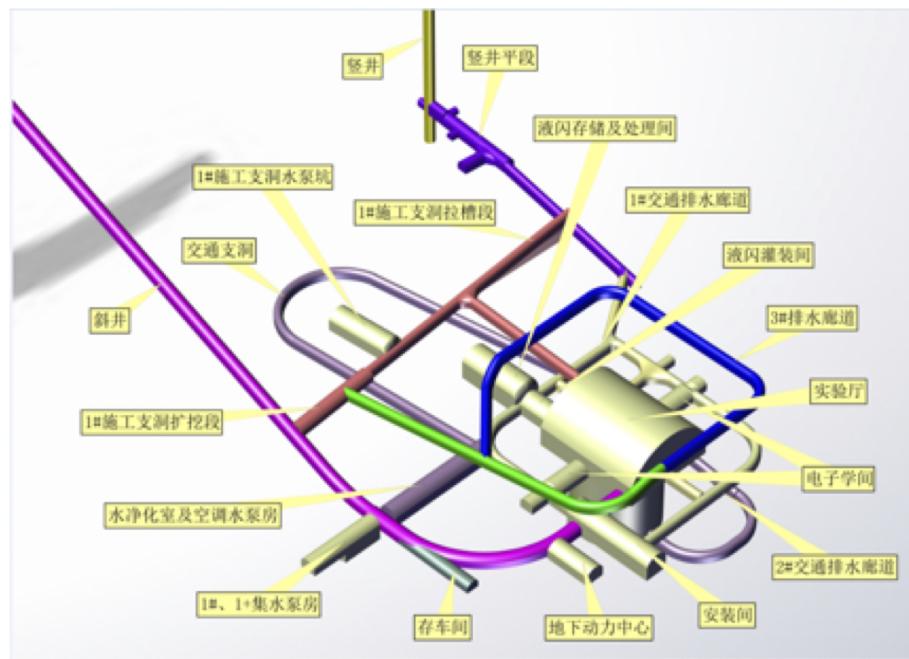


77 institutions, from 17 countries, more than 600 scientists



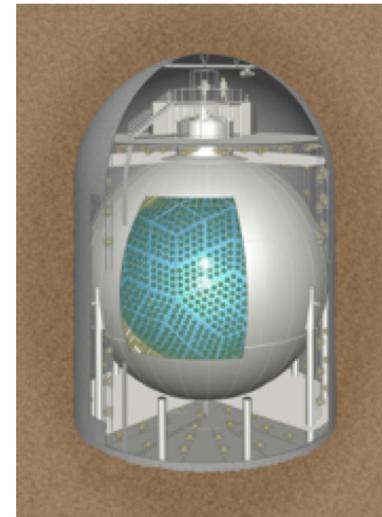


~ 600m vertical shaft, ~1300m sloped tunnel
All blasting completed on Dec. 30, 2020,
water pool construction on the way



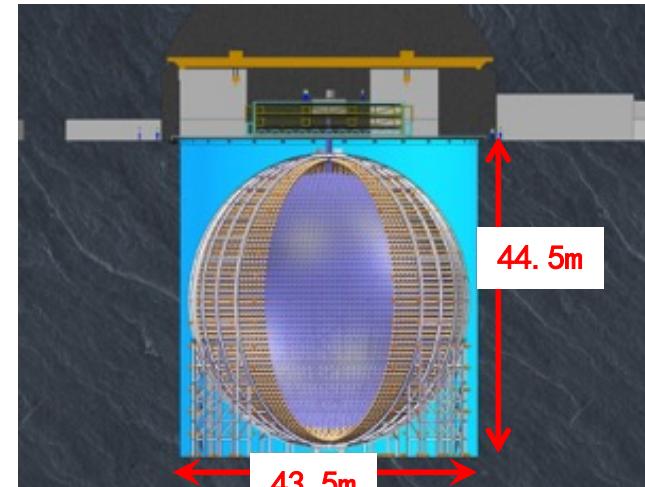
The JUNO detector

- Largest LS detector
 - × 20 KamLAND,
 - × 40 Borexino
- Highest light yield
 - × 2 Borexino,
 - × 5 KamLAND



KAMLAND

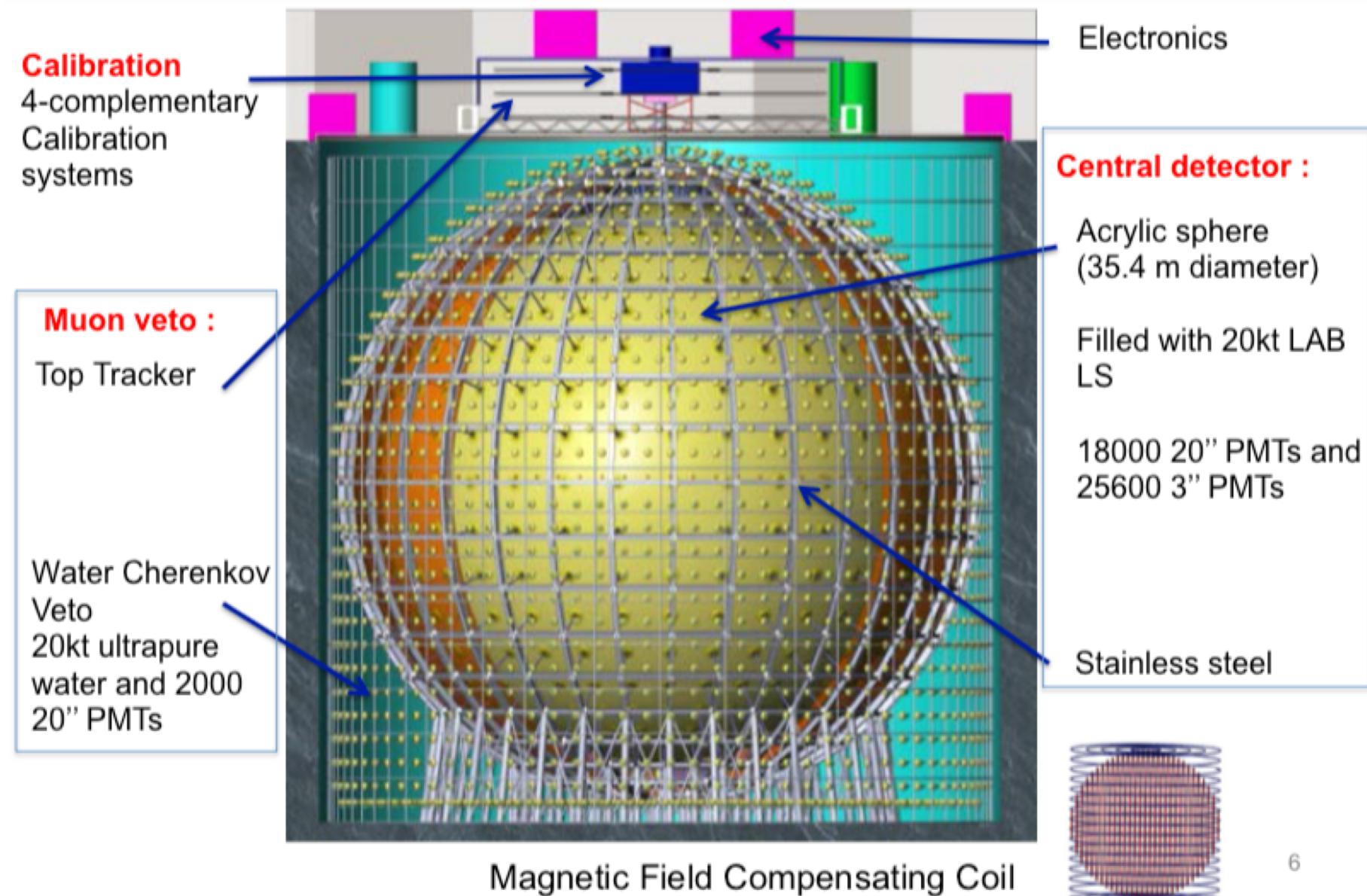
An unprecedented LS detector !



JUNO

Energy Resolution	6% @ 1MeV	3% @ 1MeV
LS mass	~1 kt	20 kt
LS Attenuation/Diameter	15m / 16m	>20m / 35m
Photocathode Coverage	32%	75%
QE x CE	25% x 60%	40% x 60%
Photon collection	250 p.e./MeV	1200 p.e./MeV

JUNO detector main components



6

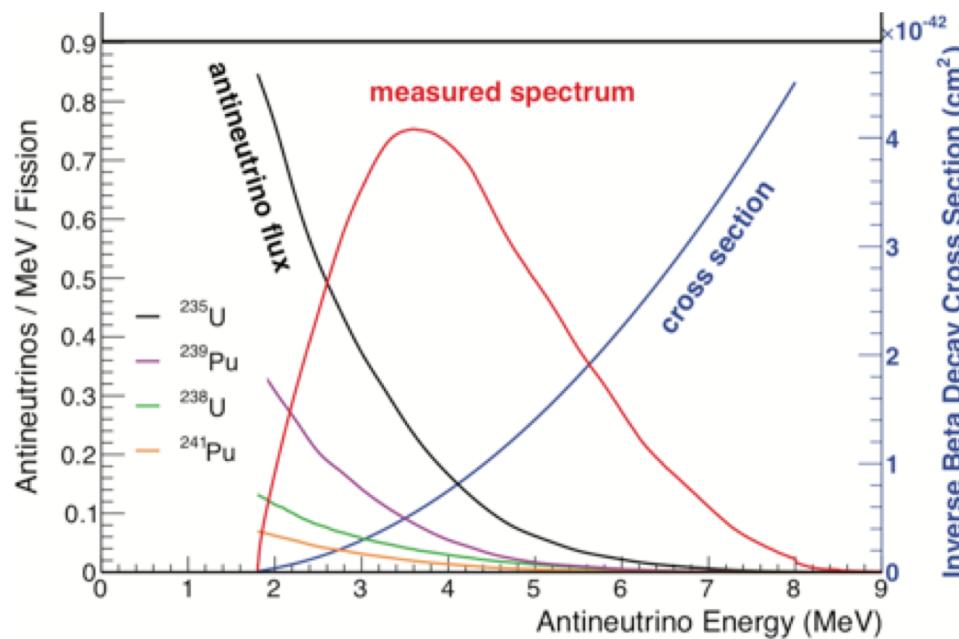
Physics at JUNO

Neutrinos are observed via inverse beta decay

$$\bar{\nu}_e + p \rightarrow e^+ + n$$

$$\tau \approx 200\mu s \quad n + p \rightarrow d + \gamma$$

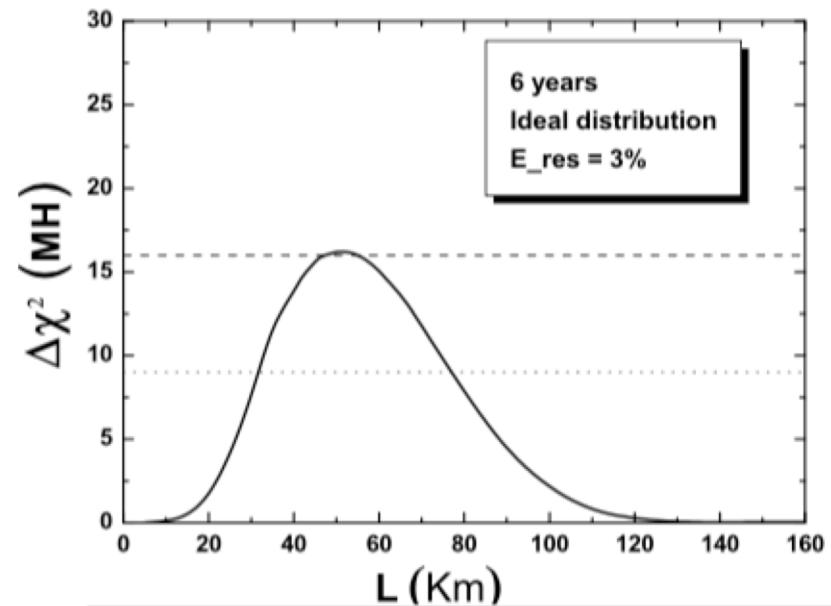
Very clean signature
E range: 2 till 8 MeV



JUNO Physics Book: arXiv: 1507.05613

	Current	JUNO
Δm^2_{12}	3%	0.6%
Δm^2_{23}	3%	0.6%
$\sin^2 \theta_{12}$	4%	0.7%
$\sin^2 \theta_{23}$	10%	N/A
$\sin^2 \theta_{13}$	DYB	

Sensitivity for 100 k IBDs
(20 kton \times 35 GW \times 6 years).



JUNO Neutrino energy spectrum

The Pee survival probability in vacuum

$$P_{ee}(L/E) = 1 - P_{21} - P_{31} - P_{32}$$

$$P_{21} = \cos^4(\theta_{13}) \sin^2(2\theta_{12}) \sin^2(\Delta_{21})$$

$$P_{31} = \cos^2(\theta_{12}) \sin^2(2\theta_{13}) \sin^2(\Delta_{31})$$

$$P_{32} = \sin^2(\theta_{12}) \sin^2(2\theta_{13}) \sin^2(\Delta_{32})$$

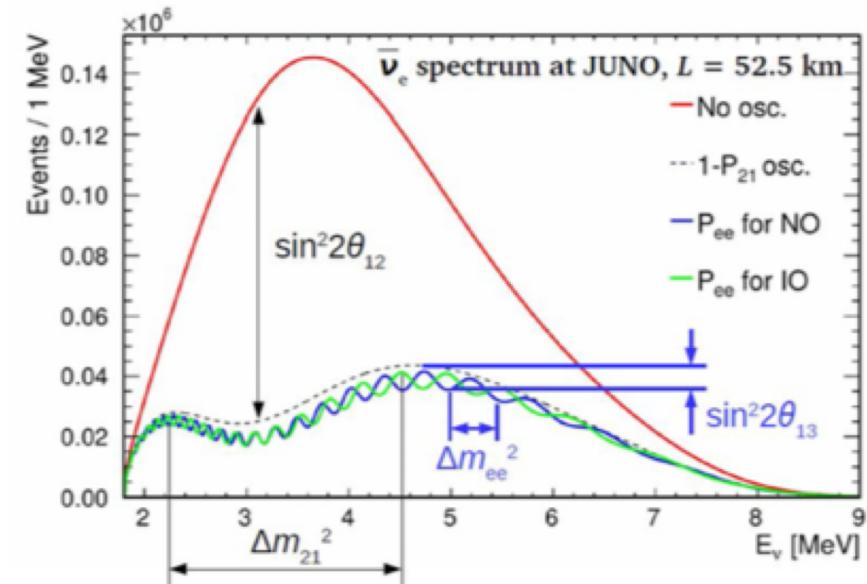
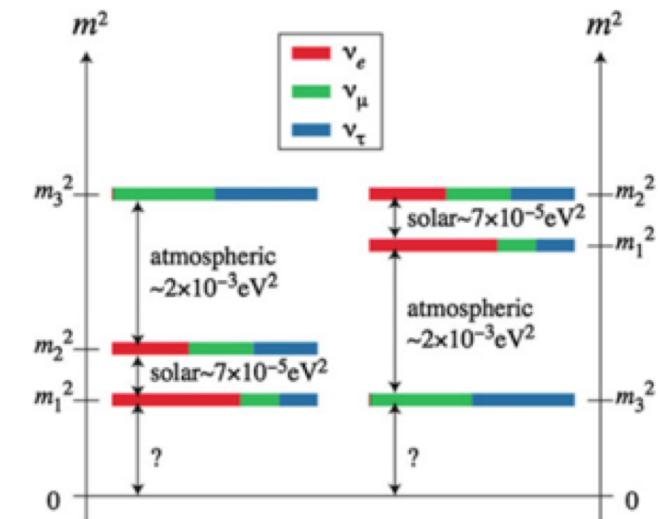
$$\Delta_{ij} \equiv \frac{\Delta m_{ij}^2 L}{4E_\nu}, \quad (\Delta m_{ij}^2 \equiv m_i^2 - m_j^2)$$

$$\Delta m_{31}^2 = \Delta m_{32}^2 + \Delta m_{21}^2$$

$$\text{NH : } |\Delta m_{31}^2| = |\Delta m_{32}^2| + |\Delta m_{21}^2|$$

$$\text{IH : } |\Delta m_{31}^2| = |\Delta m_{32}^2| - |\Delta m_{21}^2|$$

$$+ \text{NH} \quad \pm \frac{\sin^2 \theta_{12}}{2} \sin^2 2\theta_{13} \sin (2\Delta_{21}) \sin (2|\Delta_{31}|)$$



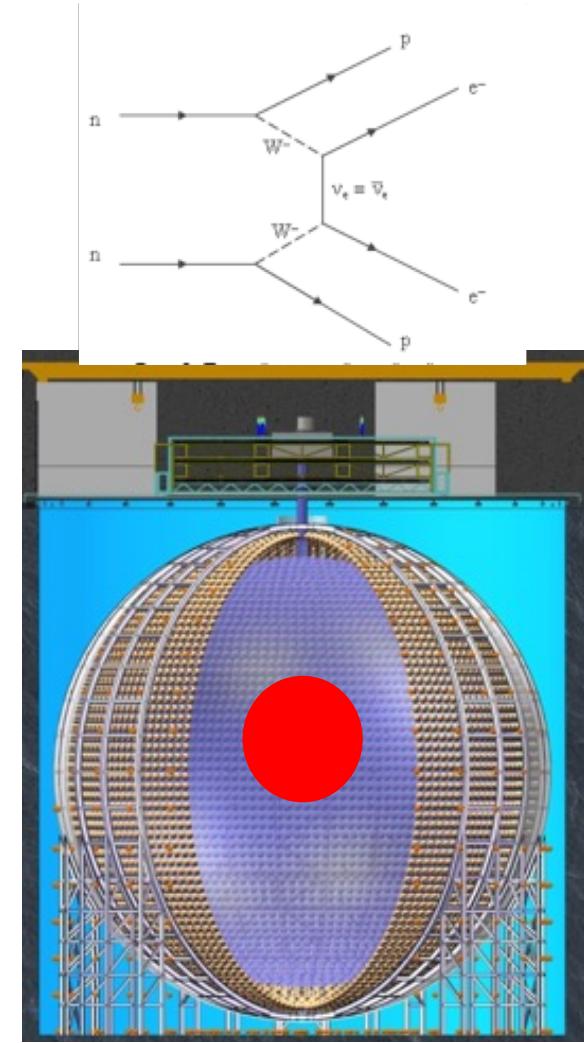
Physics at JUNO

- ❖ Mass Hierarchy: 4σ in 6 yrs, independent of the CP phase
- ❖ Neutrino oscillation parameters: precision $< 1\%$, and the test of unitarity of the mixing matrix
- ❖ Burst supernova neutrino: ~ 5000 events@10 kpc
- ❖ Discovery of diffused supernova neutrinos if $\langle E \rangle > 15$ MeV
- ❖ Geoneutrinos: 400 evts/year determine U/Th ratio of the earth
- ❖ Solar neutrinos
- ❖ Proton decays: $p \rightarrow K^- + \bar{\nu}$

JUNO - $0\nu\beta\beta$

- In ten years from now, oscillation will be mostly understood.
- Onbb decay will be the next major breakthrough
- The best limits on the lifetime of $0\nu\beta\beta$ reach 10^{25} years ($m_{\nu_{\text{majorana}}} < 0.4$ eV)
- $0\nu\beta\beta : {}^{136}_{54}\text{Xe} \rightarrow {}^{136}_{56}\text{Ba}$
 $Q(M[{}^{136}\text{Xe}] - M[{}^{136}\text{Ba}]) = 2457.83(37)$ keV
- JUNO has the potential to reach a sensitivity (at 90% C. L.) to $T_{1/2}^{0\nu\beta\beta}$ of 1.8×10^{28} yr (5.6×10^{27} yr) with ~50 tons (5 tons) of fiducial ${}^{136}\text{Xe}$ and 5 years exposure.

	Nuclides	Mass (tons)	$\langle m_{\beta\beta} \rangle, \text{meV}$
KamLAND-Zen	${}^{136}\text{Xe}$	1	61-165
EXO	${}^{136}\text{Xe}$	0.2	93-286
nEXO	${}^{136}\text{Xe}$	5	7-22
GERDA/Majorana	${}^{76}\text{Ge}$	1	10-40
SNO+	${}^{130}\text{Te}$	8	19-46
JUNO- $\beta\beta$	${}^{136}\text{Xe}$	50	4-12
	${}^{130}\text{Te}$	100-200	2-6



Insert a balloon filled with ${}^{136}\text{Xe}$ -loaded LS(or ${}^{130}\text{Te}$) into the JUNO detector

Zhao et al., arXiv: 1610.07143,
CPC 41 (2017) 5

JUNO TAO - Taishan Antineutrino Observatory

- TAO will measure the anti-neutrino spectrum at % level, to provide:

- a model-independent reference spectrum for JUNO
- a benchmark for investigation of the nuclear database

arXiv:2005.08745v1

- Start to operate 2022 year.

Detector concept :

2.8t Gd-loaded LS @ -50°C + SiPM (10m^2 | $\epsilon > 50\%$)

| 4500 e/MeV)

700k/year @ 30m from Taishan

20x JUNO 6-year data in 3 year

Energy resolution: $1.5\%/\sqrt{E}$

Status:

Design and R&D on the way:

- LS works in -50°C
- SiPM & its readout electronics
- Mechanical design
- Measured onsite muon/neutron flux
- prototype a low temperature LS detector

